# Original paper Geochemical signature of Eocene Kuh-e Dom shoshonitic dikes in NE Ardestan, Central Iran: implications for melt evolution and tectonic setting

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The Late Eocene Kuh-e Dom composite intrusion forms a segment of the Urumieh-Dokhtar magmatic arc, which recorded syn- to post-collisional magmatism during the Alpine–Himalayan orogeny in central Iran. Numerous acid and intermediate–basic dikes intrude the composite intrusive complex of the arc segment and its host-rock assemblage.

The silicic dikes of porphyric microgranite, porphyric microgranodiorite and aplite consist of quartz, K-feldspar, plagioclase (albite), biotite and rare amphibole. The dikes are of subaluminous composition with shoshonitic affinity. Trace-element patterns exhibit pronounced negative anomalies of Nb, Ta, Ti, P and Sr together with positive anomalies of Cs, Th, U and La suggesting partial melting of a quartzo-feldspathic crustal source.

The intermediate–basic dikes with phonolite, basanite and trachyandesite chemical compositions typically contain pyroxene (diopside–augite) and plagioclase phenocrysts ( $An_{30-60}$  and  $An_{98}$ ), calcic amphiboles (magnesiohornblende–magnesiohastingsite), magnesian biotites and alkali-feldspars ( $Or_{95}$ ). The rocks show shoshonitic geochemical affinities. Low Ba/Rb ratios and high Rb/Sr ratios suggest that the primary dike melt originated by partial melting of a phlogopite-bearing lithospheric mantle, whereas LILE and LREE enrichment along with low Nb/Zr and Hf/Sm ratios and high Ba/Nb ratios imply that these rocks formed at a convergent continental margin.

The acidic dikes were emplaced in a transitional syn-collisional subduction setting whereas the intermediate-basic dikes have developed from remnant melt batches after cessation of active subduction, mostly in the post-collisional setting.

Keywords: shoshonite dikes, collisional magmatism, mineral chemistry, Kuh-e Dom, petrogenesis, Iran Received: 29 March 2012; accepted: 26 November 2012; handling editor: E. Jelínek

## 1. Introduction

The Kuh-e Dom intrusion is a part of the Urumieh-Dokhtar magmatic arc, a 5-25 km wide and ~2000 km long northwest-southeast trending magmatic belt in central Iran. It is aligned parallel with the suture of the collisional Zagros fold-thrust belt and reflects the subduction of the Neotethys as Africa moved northward relative to Eurasia. Igneous activity of this magmatic arc developed from the beginning of subduction in Jurassic (Alavi 1980) and reached a climax in the Middle Eocene (Berberian & King 1981) with geochemical and petrological features similar to those of Andeantype magmatism (Berberian et al. 1982). The Neotethys is believed to have closed during Late Oligocene to Miocene (Mohajjel et al. 2003; Shahabpour 2007), initiating the collision between Arabia and Eurasia. Subduction-related magmatism in central Iran ranges from Cretaceous to Recent, but is dominated in volume by 50-35 Ma old intermediate to silicic calc-alkaline and locally shoshonitic volcanic and intrusive rocks (Berberian and King 1981; Berberian et al. 1982; Alavi and Mahdavi 1994).

The various dikes cutting the Urumieh-Dokhtar magmatic belt can be divided into two main groups: (1) silicic (porphyric microgranodiorites, porphyric microgranites and aplites), and (2) intermediate–basic (phonolite, basanite and trachyandesite). Both types exhibit shoshonitic nature which is inferred to represent a transitional stage between calc-alkaline and alkaline magmas (Liégeois et al. 1998). The dikes are characterized by high K<sub>2</sub>O/Na<sub>2</sub>O ratios (> 0.5), high Na<sub>2</sub>O + K<sub>2</sub>O (> 5 wt. %), LILE (e.g., K, Rb, Sr and Ba), and generally low TiO<sub>2</sub> contents (< 1.3 wt. %) similar to typical shoshonitic series (Joplin 1968; Morrison 1980).

Joplin (1965) suggested that shoshonitic rocks generally show SiO<sub>2</sub> less than 70 wt. % (from basaltic to trachytic rocks). On the other hand, Keller (1974) suggested that shoshonites should also include more silicic rocks such as dacite and rhyolite. Peccerillo and Taylor (1976) only regarded types with SiO<sub>2</sub> < 63 wt. % as shoshonites. Morrison (1980), proposed shoshonites to be subdivided

into basic (SiO<sub>2</sub> < 53 wt. %), intermediate (SiO<sub>2</sub> = 53–63 wt. %) and silicic (SiO<sub>2</sub> > 63 wt. %).

It seems that the chemical variability of shoshonitic rocks is not completely understood in terms of source characteristics and geodynamic setting. Morrison (1980) suggested that shoshonites are restricted to subductionrelated tectonic settings, where they are generated during late-stage of arc evolution, typically after the low-K tholeiites and calc-alkaline rock series. Similarly, Müller and Groves (1997) concluded that shoshonites should be regarded as products of magmatic activity above subduction zones in continental arcs, post-collisional arcs, or in oceanic-arc settings where potassic rocks are related to the melting of phlogopite-bearing wedge.

The aim of this study is to contribute to the understanding of the acidic and intermediate-basic dikes in the Kuh-e Dom intrusion of the Urumieh-Dokhtar magmatic belt. More specifically, this contribution deals with the petrogenesis and tectonic setting of the dikes, by evaluating field, petrographic, whole-rock geochemical and mineral compositional data.

## 2. Analytical methods

The petrological variation and the crosscutting relations of the dikes were observed during field analysis. Both petrographic and geochemical analyses were conducted using the most representative and fresh samples from various lithologies in the studied intrusions. The thirteen samples were selected for whole-rock geochemical analysis and 90 points for microprobe analysis.

The mineral analyses were performed in the electron microprobe Laboratory at Naruto University, Japan (Jeol JXA-8800R) equipped by WDS detector using 15 kV, 20 nA and 20 s counting time and University of Oklahoma, Norman, USA (Cameca SX50) with five asynchronous wavelength-dispersive spectrometers and PGT PRISM 2000 energy-dispersive X-ray analyzer. Petrographic characterization was performed by backscattered electron imaging coupled with energy-dispersive X-ray analysis using 20 kV acceleration voltage and 20 nA sample current. Chemical microanalysis was carried out by wavelength-dispersive spectrometry using 20 kV acceleration voltage, 20 nA beam current (measured at the Faraday cup), and 2 µm spot size. Matrix corrections employed the PAP algorithm (Pouchou and Pichoir 1985), with oxygen content calculated by stoichiometry. Standard materials were natural crystalline solids for all elements, for which NIST glass K309 was used. Counting times were 30 s on peak for all elements, yielding minimum levels of detection (calculated at  $3\sigma$  above mean background) in the range of 0.01-0.02 wt. % of the oxides.

Major elements were analyzed by X-ray fluorescence (XRF) apparatus Rigaku RIX 2000 with a Rh tube

hosted at Naruto University in Japan. Glass beads from finely ground samples were prepared with a sampleto-flux ( $\text{Li}_2\text{B}_4\text{O}_7$ ) ratio of 1:10 and the method yielded analytical errors < 1 %. Some major- and trace-element compositions, including REE, were determined at ALS Chemex, Vancouver, Canada by ICP-AES and ICP-MS techniques. Major elements were analyzed by ICP-AES (detection limit ~ 0.01 %). Trace-element analyses were carried out by ICP-MS (detection limit of 0.01–0.1 ppm), following Li metaborate fusion and HNO<sub>3</sub> total digestion. Details of the chemical procedures are found at *http://www.alschemex.com*.

## 3. Geological setting

#### 3.1. Kuh-e Dom intrusion

This intrusion is an arc segment of the Urumieh-Dokhtar magmatic belt. It includes gabbro, diorite, quartz diorite, monzodiorite, monzonite, quartz monzonite, quartz monzodiorite, monzogranite and granodiorite, and – based on K-Ar dating studies – was emplaced in Late Eocene (Technoexport 1981). The intermediate-basic units form a discontinuous band surrounding the northern, southern and eastern margins. The contact zone where I-type, high-K calc-alkaline, metaluminous Kuh-e Dom plutons (Sarjoughian 2007) intruded Paleozoic schists, Cretaceous limestones and early Eocene volcanic rocks (Fig. 1) is marked by a thermal metamorphic aureole with skarn and hornfels. Crosscutting relations show that the intermediate-basic are relatively younger than the silicic igneous rocks. The whole-rock geochemical signature indicates that the parental magma may have been derived from a mantle and lower crust sources carrying a subductionrelated imprint. Underplating of already thickened continental crust by mantle-derived hot basic magma was proposed to have resulted in dehydration melting of the lower crust and mixing between the contrasting magmas (Sarjoughian 2012).

#### 3.2. Silicic dikes

The silicic dikes (porphyric microgranites, porphyric microgranodiorites and aplites) are oriented W–E to NW–SE, forming c. 10 % of igneous exposures. They tend to be cream or light gray in color. The dikes intruded silicic and intermediate–basic units of the intrusion and the host rocks, especially in its southern and southeastern parts. The dikes generally show subhorizontal orientations, although sometimes they clearly cross cut each other. Although the thicknesses of silicic dikes vary between 0.5 and 20 meters, most are 5 to 6 meters wide. The studied



(b)



Fig. 1a Generalized tectonic map of Iran: UDMA = Urumieh-Dokhtar Magmatic Arc, SSZ = Sanandaj-Sirjan Zone, MZT = Main Zagros Thrust. b – Simplified geological map of Kuh-e Dom intrusion adopted from Technoexport (1981), slightly modified.



Fig. 2 Photomicrographs (XPL) of porphyritic microgranite (H3) with corroded quartz phenocrysts surrounded by reaction rims (a) and porphyritic microgranodiorite (F12) with sieve texture in plagioclase (b).

dikes are characterized mostly by porphyric texture and minor chloritization. Contacts between the dikes and plutonic host rocks are sharp and planar, implying that the pluton was rigid at the time of dike emplacement. The dikes with chilled margins indicate emplacement at shallow depths after the coarse-grained granite had solidified.

#### 3.3. Intermediate-basic dikes

The dark green or brown intermediate–basic dikes cutting the Kuh-e Dom intrusion and host rocks are oriented W–E and occur less frequently than their silicic counterpants. They are chemically classified as phonolite, basanite and trachyandesite with partly aphanitic and porphyric texture, fine to medium-grained. These dikes are c. 1 meter thick and have no obvious chilled margins. Most intermediate–basic dikes show hydrothermal alteration with local calcitic, chloritic, epidotic and spilititic veins.

# 4. Petrography

#### 4.1. Silicic dikes

The rock-forming minerals include quartz, alkali-feldspar, plagioclase, biotite and rarely hornblende, all of which occur both as phenocrysts and in the groundmass. The phenocrysts are present either as isolated phenocrysts or as glomerocrysts.

Alkali-feldspar and plagioclase usually dominate both the phenocryst assemblage and the groundmass. However, some porphyric microgranite dikes contain more than 90 % of quartz phenocrysts with low contents of alkali-feldspar, plagioclase and biotite phenocrysts. The coarse quartz grains often show corroded margins; some contain dissolution cavities in central parts. Quartz phenocrysts are surrounded by reaction rims of fine-grained feldspars (Fig. 2a), which indicate disequilibrium of the quartz crystals with the surrounding magma, caused by changes in melt pressure or chemical composition (Vernon 2004).

Alkali-feldspars occur more frequently in porphyric microgranite than in the porphyric microgranodiorite. The latter shows large amounts of plagioclase phenocrysts and or as microlites. The phenocrysts are often euhedral indicating early crystal growth, whereas the plagioclase microlites form the groundmass of these rocks. Some large crystals have dissolved margins caused by reactions with the surrounding magma. Plagioclase phenocrysts show polysynthetic twinning and/or oscillatory zoning indicating a slow equilibration relative to the crystallization rate (Vernon 2004).

Plagioclases enclose various mineral inclusions and show poikilitic and, rarely, sieve-textures (Fig. 2b). The sieve texture is apparently caused by changes in crystal growth conditions causing dissolution of parts of the crystal. As the crystals continued to grow, portions of melt were surrounded by plagioclase (Vernon 2004). The mafic minerals of these dikes are common euhedral, fineto medium-grained biotites and rare euhedral amphiboles. The phenocrysts are embedded in a fine-grained groundmass formed of quartz and feldspar, and some biotite. Apatite and zircon are accessory minerals. Some dikes show sericitization and kaolinization of alkali feldspars, saussuritization and carbonatization of plagioclase, and chloritization of biotites and hornblendes.

The aplites consist of abundant anhedral quartz, alkalifeldspar and rarely albitic plagioclase and biotite. The alkalifeldspars tend to be kaolinized and show perthitic texture. The plagioclases often show polysynthetic twinning and are slightly sericitized. Biotite as anhedral crystals is the only ferromagnesian mineral; it is commonly altered to chlorite.



Fig. 3 Photomicrographs (XPL) of intermediate-basic dikes.  $\mathbf{a}$  - Corroded plagioclase with discontinuous zoning (S10).  $\mathbf{b}$  - Poikilitic texture of clinopyroxene (S35).

#### 4.2. Intermediate-basic dikes

The intermediate-basic dikes usually show porphyric and, sometimes, microlitic textures. They are composed of 60-70 % pyroxene and plagioclase phenocrysts set in groundmass of plagioclase microlites with fine-grained pyroxene, plagioclase, amphibole, biotite and alkali-feldspar. Basanites and trachyandesite are fine-to medium-grained and differ from phonolites by the presence of abundant pyroxene, lower feldspars and higher mafic minerals contents. Plagioclase is the dominant phenocryst phase, forming mainly subhedral and lath-shaped crystals (0.3 mm to several mm across). Frequently they show corroded margins (Fig. 3a) and enclose many inclusions of clinopyroxene and opaque minerals. The plagioclase crystals show polysynthetic twinning, oscillatory and/or discontinuous zoning, as well as sieve textures. Some grains are altered to sericite and clay minerals.

Subhedral clinopyroxene is the main mafic phase in the dikes. This mineral encloses plagioclase and mafic mineral inclusions, forming poikilitic textures (Fig. 3b). It suffered intense alteration; some clinopyroxenes are partially replaced by actinolitic amphibole, others are completely uralitized and carbonatized.

Rare euhedral prismatic amphibole crystals of green color are associated with biotite and slightly altered to chlorite. Biotite is generally subhedral, pleochroic from reddish brown to brown and often altered to titanite and chlorite. K-feldspar mostly occupies the interstices between plagioclase laths and other minerals, suggesting late crystallization. Some of the K-feldspar rims the plagioclase phenocrysts. Acicular apatite, sphene and opaque minerals are common accessory phases. Some of the dikes were altered to spilite and are composed now of oligoclase, together with chlorite, epidote, calcite, and actinolite.

## 5. Chemical composition of minerals

The results of electron-microprobe analysis are shown in Tables 1-5. Mineral recalculations were done by the MinPet Geological Software version 2.02 (Richard 1995). Structural formulae of feldspars were calculated on the basis of 32 oxygens. Pyroxenes were classified by using the total numbers of specified cations at the M sites on the basis of 6 O atoms. The Fe<sup>3+</sup> estimates were made using the general equation of Droop (1987) assuming 4 cations per formula unit. Amphibole analyses were recalculated on anhydrous basis to cations per 23 oxygens. Ferrous and ferric iron were assessed by the charge balance method described by Robinson et al. (1981), using the average between the 15-NK and 13-CNK methods after applying the appropriate IMA site distribution rules (Leake et al. 1997). Biotite formulae were recalculated based on 24 O atoms and no OH groups; the oxidation state of iron was assessed by the Mica+ software (Yavuz 2003).

In the ternary Ab–Or–An diagram (Deer et al. 1992), plagioclase compositions in the silicic dikes are albite  $(An_{3-9})$  and in the intermediate–basic dikes andesine, labradorite, and anorthite  $(An_{30-56} \text{ and } An_{98}; \text{ Fig. 4})$ . Alkali-feldspar in intermediate–basic dikes corresponds to orthoclase with Or content between 60 and 95 % (Fig. 4).

In the intermediate dikes, biotites with high TiO<sub>2</sub> contents are primary (Nachit et al. 2005) and classify (Rieder et al. 1998) as biotite close to phlogopite end member ( $X_{Mg} = Mg/(Mg + Fe) \sim 0.63$ ; Fig. 5a). The alumina saturation index (ASI) of biotites [AI<sup>T</sup>/(Ca + Na + K)] is low (1.28) and reflects lower alumina activity in the magma (Zen 1988). Based on FeO and Al<sub>2</sub>O<sub>3</sub> contents, all analyses plot into the calc-alkaline field (Fig. 5b; Abdel-Rahman 1994). Mg-rich micas with low Fe<sup>2+</sup>/(Mg+Fe<sup>2+</sup>) ratios plot in the shoshonitic field (Fig. 5c; Jiang et al. 2002), indicative of high  $fO_2$  conditions of the magmas during crystallization.

Tab.	l Electron-microprobe	analyses of plagiocla	se in Kuh-e Dom dik	kes (wt. % and	apfu calculated o	n the basis of 32 O).
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Sample	F117-21	F117-22	F117-41	F117-42	F117-43	F117-44	\$28-1	\$28-3
Location	dik A	dik A	dik A	dik A	dik A	dik A	dik B	dik B
SiO.	67.55	68.33	66.83	68.98	64.32	68.15	53.91	56.00
TiO.	0.00	0.01	0.00	0.00	0.00	0.03	0.03	0.16
ALO.	20.03	19.65	20.02	19.51	20.75	19.91	28.84	27.44
FeO	0.12	0.09	0.55	0.05	0.49	0.05	0.58	0.50
MnO	0.01	0.00	0.00	0.01	0.00	0.00	0.00	0.01
ΜσΟ	0.04	0.02	0.09	0.00	0.12	0.01	0.00	0.05
BaO	0.09	0.11	0.04	0.09	0.12	0.18	0.05	0.04
CaO	0.12	0.12	0.06	0.05	0.09	0.19	10.94	9.43
Na O	11.55	11 47	11.36	11.76	10.24	11.83	5 19	6.06
K O	0.27	0.20	0.46	0.12	1.63	0.11	0.24	0.00
SrO	0.027	0.03	0.03	0.01	0.01	*	0.15	0.10
Total	99.80	100.03	99.44	100.59	97.75	100.46	99.97	100.00
Si	11.86	11.95	11.81	11.99	11.63	11.89	9.78	10.10
Al	4.14	4.05	4.17	3.99	4 42	4 09	6.16	5.83
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02
Fe <sup>2+</sup>	0.02	0.01	0.08	0.01	0.07	0.01	0.09	0.08
Mn	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mg	0.00	0.00	0.02	0.00	0.00	0.00	0.00	0.00
Ra	0.01	0.01	0.02	0.00	0.03	0.00	0.01	0.00
Ca	0.02	0.02	0.00	0.01	0.01	0.04	2.13	1.82
Na	3.93	3.89	3.89	3.96	3 59	4 00	1.83	2.12
K	0.06	0.05	0.10	0.03	0.38	0.02	0.06	0.05
Ab	97.90	98.30	97.10	99.10	90.10	98.50	45.50	53.10
An	0.60	0.60	0.30	0.30	0.40	0.90	43.30 53.10	45 70
Or	1.50	1.10	2.60	0.30	9.40	0.50	1 40	1 20
01	1.50	1.10	2.00	0.70	9.40	0.00	1.40	1.20
Samula	529 14	529 15	5282	529.21	528.22	520.2	S29 21	528.22
Lagation	526-14	526-13	SZ6Z	526-21	520-22	526-5	526-51	320-32
<u>Elocation</u>	52.40	58.40	52.77	54.01	52.05	52.11	52.76	54.15
310 <sub>2</sub>	0.07	0.02	0.07	0.08	0.00	0.11	52.70	0.06
	0.07	0.03	0.07	0.08	0.09	0.11	0.00	0.08
$AI_2O_3$	29.03	0.30	26.95	28.38	28.71	29.00	26.79	20.70
FeO MrO	0.43	0.39	0.45	0.44	0.34	0.70	1.70	0.01
MaQ	0.00	0.01	0.01	0.01	0.03	0.02	0.00	0.01
MgO	0.00	0.13	0.09	0.07	0.10	0.03	0.07	0.08
GaO	11.82	6.84	0.11	10.01	10.00	11.44	0.11	10.70
CaO No O	11.82	7.50	5.06	5 17	5 20	11.44	11.38	5 26
K O	4.84	0.87	0.24	0.25	0.25	4.65	4.89	0.10
K <sub>2</sub> 0	0.17	0.87	0.24	0.23	0.23	0.17	0.21	0.19
Total	100.62	0.11	0.13	0.18	100.07	0.13	100.21	100.66
<u>s</u> ;	9.64	10.54	99.94	99.80	0.70	99.02	0.64	0.70
A1	9.04 6.21	5 29	9.70 6.10	5.82	6.12	5.03	5.04	5.75
AI T:	0.51	3.38	0.19	0.12	0.13	0.23	0.19	0.11
11 Eo <sup>2+</sup>	0.01	0.00	0.01	0.01	0.01	0.02	0.01	0.01
re	0.07	0.00	0.07	0.07	0.08	0.11	0.27	0.18
Ma	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00
IVI g	0.00	0.04	0.02	0.02	0.03	0.01	0.02	0.02
Ба	0.01	0.01	0.01	0.01	0.00	0.00	0.01	0.01
Ca No	2.29	1.32	2.15	2.13	2.14	2.24	2.23	2.09
ina K	1.70	2.62	1.78	1.82	1.83	1.72	1./3	1.84
<u>K</u>	0.04	0.20	0.06	0.06	0.06	0.04	0.05	0.04
AD	42.20	21.00	44./U	45.50	45.50	43.00	43.20	40.40
An	56.90	31.90	53.90	53.10	53.10	56.00	55.60	52.50
Or	1.00	4.80	1.40	1.40	1.40	1.00	1.20	1.10

Sample	\$28-33	S28-34	\$28-35	S28-4	S28-5	S28-6	\$35-51	\$35-52
Location	dik B	dik B	dik B	dik B	dik B	dik B	dik B	dik B
SiO <sub>2</sub>	53.06	54.46	53.40	54.78	53.13	60.03	46.16	45.93
TiO,	0.08	0.06	0.09	0.09	0.08	0.04	0.02	0.02
Al <sub>2</sub> O <sub>3</sub>	28.65	28.20	29.14	28.17	29.09	24.77	24.91	24.96
FeO	1.09	0.29	0.42	0.63	0.42	0.32	0.82	0.63
MnO	0.00	0.00	0.01	0.01	0.00	0.00	0.02	0.04
MgO	0.18	0.01	0.06	0.09	0.12	0.04	0.03	0.00
BaO	0.08	0.11	0.12	0.10	0.04	0.07	27.66	27.13
CaO	11.34	10.64	11.49	10.47	11.46	6.34	0.06	0.06
Na <sub>2</sub> O	4.83	5.35	4.87	5.51	4.81	7.86	0.09	0.24
K <sub>2</sub> O	0.19	0.22	0.25	0.27	0.22	0.32	0.07	0.09
SrO	0.14	0.16	0.17	0.16	0.19	0.09	0.07	0.34
Total	99.64	99.50	100.02	100.28	99.56	99.88	99.91	99.44
Si	9.70	9.91	9.70	9.91	9.69	10.74	8.85	8.85
Al	6.17	6.04	6.23	6.00	6.25	5.22	5.62	5.66
Ti	0.01	0.01	0.01	0.01	0.01	0.01	0.00	0.00
Fe <sup>2+</sup>	0.17	0.04	0.06	0.10	0.06	0.05	0.13	0.10
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01
Mg	0.05	0.00	0.02	0.02	0.03	0.01	0.01	0.00
Ba	0.01	0.01	0.01	0.01	0.00	0.01	0.01	0.02
Ca	2.22	2.07	2.24	2.03	2.24	1.22	5.68	5.60
Na	1.71	1.89	1.72	1.93	1.70	2.73	0.03	0.04
K	0.04	0.05	0.06	0.06	0.05	0.07	0.02	0.08
Ab	43.00	47.00	42.80	48.00	42.60	67.90	0.50	0.60
An	55.80	51.70	55.80	50.40	56.10	30.30	99.20	97.90
Or	1.10	1.30	1.40	1.50	1.30	1.80	0.30	1.50

Sample	S35-53	\$35-51	S35-61	S35-62	S35-63	S35-64	S35-65
Location	dik B	dik B	dik B	dik B	dik B	dik B	dik B
SiO <sub>2</sub>	46.02	46.28	45.95	45.76	45.68	46.19	45.39
TiO <sub>2</sub>	0.03	0.01	0.01	0.13	0.04	0.01	0.08
Al <sub>2</sub> O <sub>3</sub>	24.91	24.62	24.87	24.91	25.03	25.16	24.68
FeO	0.65	1.28	0.93	0.80	1.31	0.59	0.78
MnO	0.02	0.04	0.04	0.02	0.03	0.04	0.05
MgO	0.00	0.00	0.02	0.03	0.18	0.00	0.06
BaO	27.40	26.83	27.15	27.77	27.21	27.60	27.68
CaO	0.09	0.05	0.01	0.01	0.05	0.00	0.02
Na <sub>2</sub> O	0.10	0.17	0.11	0.00	0.10	0.09	0.15
K <sub>2</sub> O	0.12	0.07	0.09	0.06	0.05	0.06	0.17
SrO	0.13	0.40	0.28	0.06	0.05	0.13	0.16
Total	99.49	99.73	99.47	99.55	99.74	99.87	99.21
Si	8.86	8.90	8.85	8.80	8.79	8.84	8.79
Al	5.65	5.58	5.64	5.64	5.67	5.67	5.63
Ti	0.01	0.00	0.00	0.02	0.01	0.00	0.01
Fe <sup>2+</sup>	0.11	0.21	0.15	0.13	0.21	0.09	0.13
Mn	0.00	0.01	0.01	0.00	0.01	0.01	0.01
Mg	0.00	0.00	0.01	0.01	0.05	0.00	0.02
Ba	0.01	0.01	0.01	0.00	0.01	0.01	0.01
Ca	5.65	5.53	5.60	5.72	5.61	5.66	5.74
Na	0.04	0.02	0.04	0.02	0.02	0.02	0.06
К	0.03	0.10	0.07	0.02	0.01	0.03	0.04
Ab	0.80	0.40	0.60	0.40	0.40	0.40	1.10
An	98.70	97.90	98.20	99.30	99.40	99.10	98.30
Or	0.50	1.70	1.20	0.30	0.20	0.50	0.70

dik A: acidic dikes, dik B: basic-intermediate dikes

Tab.	2 Electron-microprobe	analyses of K-feld	par from intermediate	e-basic dikes (wt. 9	% and apfu calculated	on the basis of 32 O).
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Sample	S28-11	S28-12	S28-42	S28-43	S35-11	S35-12	S35-13	S35-14
SiO <sub>2</sub>	64.11	62.87	63.09	62.26	64.07	63.23	64.08	62.46
TiO <sub>2</sub>	0.05	0.06	0.04	0.38	0.00	0.06	0.00	0.00
Al <sub>2</sub> O <sub>3</sub>	18.52	18.06	19.43	16.28	18.50	18.58	18.63	18.32
FeO	0.22	0.16	0.34	2.84	0.15	0.70	0.17	1.63
MnO	0.00	0.02	0.01	0.06	0.00	0.00	0.01	0.01
MgO	0.01	0.03	0.05	2.23	0.03	0.44	0.06	1.10
BaO	0.44	0.43	0.41	0.42	0.61	0.62	0.33	0.67
CaO	0.10	0.08	1.30	2.45	0.09	0.42	0.05	0.18
Na <sub>2</sub> O	1.08	0.97	4.01	1.02	0.48	0.54	0.36	0.52
K <sub>2</sub> O	14.68	14.99	11.17	11.78	15.91	14.87	16.01	14.39
SrO	0.23	0.17	0.19	0.20	0.16	0.15	0.21	0.09
Total	99.44	97.84	100.04	99.92	100.00	99.61	99.91	99.37
Si	11.93	11.92	11.63	11.66	11.92	11.81	11.92	11.73
Al	4.06	4.03	4.22	3.59	4.05	4.09	4.08	4.05
Ti	0.01	0.01	0.01	0.05	0.00	0.01	0.00	0.00
$Fe^{2+}$	0.03	0.03	0.05	0.45	0.02	0.11	0.03	0.26
Mn	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00
Mg	0.00	0.01	0.01	0.62	0.01	0.12	0.02	0.31
Ba	0.03	0.03	0.03	0.03	0.04	0.05	0.02	0.05
Ca	0.02	0.02	0.26	0.49	0.02	0.08	0.01	0.04
Na	0.39	0.36	1.43	0.37	0.17	0.20	0.13	0.19
Κ	3.49	3.63	2.63	2.81	3.78	3.54	3.80	3.45
Ab	10.00	8.90	33.20	10.10	4.40	5.10	3.30	5.10
An	0.50	0.40	6.00	13.40	0.50	2.20	0.30	1.00
Or	89.50	90.70	60.80	76.60	95.20	92.70	96.40	93.90
Sample	\$35-15	\$35-1	6 5	35_61	\$35_62	\$35-63	\$35-71	\$35-72

Sample	\$35-15	\$35-16	\$35-61	\$35-62	\$35-63	\$35-71	\$35-72
SiO <sub>2</sub>	64.53	61.83	63.40	62.59	63.12	64.22	64.30
TiO <sub>2</sub>	0.03	0.00	0.00	0.07	0.00	0.01	0.03
Al <sub>2</sub> O <sub>3</sub>	18.09	18.24	17.99	19.01	18.86	18.81	18.75
FeO	0.08	2.06	0.41	1.33	0.27	0.02	0.02
MnO	0.00	0.03	0.01	0.02	0.01	0.00	0.01
MgO	0.03	1.27	0.32	0.44	0.07	0.00	0.01
BaO	0.26	0.57	0.67	0.85	0.66	0.35	0.32
CaO	0.05	0.22	0.17	0.15	0.24	0.00	0.01
Na <sub>2</sub> O	0.35	0.36	0.44	0.41	0.41	0.41	0.39
K <sub>2</sub> O	16.03	14.37	15.29	14.98	15.66	15.83	15.92
SrO	0.13	0.15	0.22	0.22	0.19	0.55	0.46
Total	99.58	99.10	98.92	100.07	99.49	100.20	100.22
Si	12.01	11.67	11.93	11.70	11.82	11.92	11.93
Al	3.97	4.06	3.99	4.19	4.16	4.11	4.10
Ti	0.00	0.00	0.00	0.01	0.00	0.00	0.00
Fe <sup>2+</sup>	0.01	0.33	0.07	0.21	0.04	0.00	0.00
Mn	0.00	0.01	0.00	0.00	0.00	0.00	0.00
Mg	0.01	0.36	0.09	0.12	0.02	0.00	0.00
Ba	0.02	0.04	0.05	0.06	0.05	0.03	0.02
Ca	0.01	0.04	0.03	0.03	0.05	0.00	0.00
Na	0.13	0.13	0.16	0.15	0.15	0.15	0.14
Κ	3.81	3.46	3.67	3.57	3.74	3.75	3.77
Ab	3.20	3.60	4.20	4.00	3.80	3.80	3.60
An	0.30	1.20	0.90	0.80	1.20	0.00	0.10
Or	96.50	95.20	95.00	95.20	95.00	96.20	96.40

Sample	S28-11	S28.2	S28-13	S28-31	S28-32	S28-41	S282	S28-43
SiO <sub>2</sub>	37.67	37.58	37.48	37.86	38.07	38.25	37.99	37.94
TiO <sub>2</sub>	4.37	4.26	4.58	4.55	4.41	4.37	4.31	4.03
Al <sub>2</sub> O <sub>3</sub>	13.26	13.24	13.29	13.37	13.28	13.43	13.36	13.09
Cr <sub>2</sub> O <sub>3</sub>	0.02	0.01	0.02	0.03	0.03	0.03	0.04	0.04
FeO	15.39	16.20	15.30	15.66	15.11	15.61	15.76	15.69
MnO	0.17	0.19	0.16	0.19	0.20	0.16	0.17	0.19
MgO	14.58	14.59	14.46	14.92	14.64	15.02	15.14	15.02
BaO	0.13	0.15	0.13	0.15	0.14	0.14	0.20	0.12
CaO	0.02	0.02	0.04	0.03	0.21	0.02	0.04	0.18
Na <sub>2</sub> O	0.07	0.04	0.05	0.10	0.07	0.11	0.08	0.02
K <sub>2</sub> O	9.67	9.54	9.43	9.45	9.43	9.43	9.58	9.57
F	0.24	0.29	0.23	0.14	0.29	0.19	0.50	0.24
Cl	0.32	0.30	0.28	0.31	0.29	0.29	0.29	0.32
Total	95.91	96.41	95.45	96.76	96.17	97.05	97.46	96.45
Si	5.91	5.89	5.90	5.88	5.94	5.92	5.89	5.93
Al <sup>IV</sup>	2.09	2.11	2.10	2.12	2.06	2.08	2.11	2.07
Al <sup>VI</sup>	0.37	0.34	0.37	0.33	0.38	0.36	0.33	0.34
Ti	0.52	0.50	0.54	0.53	0.52	0.51	0.50	0.47
Fe <sup>3+</sup>	0.65	0.78	0.70	0.82	0.58	0.74	0.78	0.66
Fe <sup>2+</sup>	1.28	1.25	1.20	1.13	1.31	1.19	1.18	1.30
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01
Mn	0.02	0.03	0.02	0.03	0.03	0.02	0.02	0.03
Mg	3.41	3.41	3.40	3.46	3.41	3.46	3.50	3.50
Ba	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01
Ca	0.00	0.00	0.01	0.01	0.04	0.00	0.01	0.03
Na	0.02	0.01	0.02	0.03	0.02	0.03	0.02	0.01
Κ	1.94	1.91	1.90	1.87	1.88	1.86	1.89	1.91
F	0.24	0.29	0.23	0.14	0.29	0.19	0.49	0.24
Cl	0.17	0.16	0.15	0.16	0.15	0.15	0.15	0.17

Tab. 3 Electron-microprobe analyses of biotite from intermediate dikes (wt. % and apfu recalculated based on 24 O).



**Fig. 4** Compositions of feldspars from the Kuh-e Dom dikes plotted into the An–Ab–Or ternary diagram (Deer et al. 1992).

Sample	S35-2C1	S35-2M1	S35-2M2	S35-2R1	S35-3C1	S35-3R1	S35-4C1	S35-4R1	S35-5C2	S35-5CR1	S35-5R3	S35-6M1
	52.69	52.35	51.55	53.14	52.89	51.84	52.82	52.96	51.65	52.40	52.79	49.02
$IO_2$	0.19	0.27	0.60	0.13	0.23	0.40	0.12	0.11	0.20	0.23	0.17	0.78
$\mathbf{M}_2\mathbf{O}_3$	0.86	1.59	3.09	0.73	0.94	2.46	0.84	0.70	1.80	1.09	0.98	4.94
eO	10.42	9.49	8.65	10.50	9.54	9.10	10.36	10.42	11.39	10.16	10.19	10.17
$r_2O_3$	0.02	0.01	0.01	0.00	0.00	0.01	0.01	0.00	0.01	0.01	0.00	0.00
InO	0.62	0.54	0.38	0.71	0.46	0.41	0.63	0.70	0.58	0.65	0.61	0.28
1gO	13.79	14.77	14.84	13.91	14.50	14.74	13.83	13.81	13.36	13.77	13.93	13.61
aO	21.20	20.64	20.68	21.25	21.17	20.91	21.31	21.30	20.37	20.94	21.10	20.51
$a_2O$	0.24	0.30	0.28	0.24	0.22	0.29	0.23	0.22	0.21	0.30	0.24	0.28
0	00.00	0.02	0.02	0.00	0.01	0.00	0.00	0.00	0.02	0.00	0.00	0.04
otal	100.03	99.98	100.10	100.61	96.66	100.16	100.15	100.22	99.59	99.55	100.01	99.63
Si	1.97	1.95	1.91	1.98	1.97	1.92	1.98	1.98	1.95	1.97	1.97	1.83
AI	0.03	0.05	0.09	0.02	0.03	0.08	0.03	0.02	0.05	0.03	0.03	0.17
$\mathrm{Fe}^{3+}$	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
IIAI	0.01	0.02	0.04	0.01	0.01	0.03	0.01	0.01	0.03	0.02	0.02	0.05
ПТі	0.01	0.01	0.02	0.00	0.01	0.01	0.00	0.00	0.01	0.01	0.01	0.02
$11 \mathrm{Fe}^{3+}$	0.02	0.04	0.04	0.02	0.02	0.05	0.02	0.02	0.03	0.02	0.02	0.10
$11 \mathrm{Fe}^{2+}$	0.19	0.11	0.09	0.19	0.16	0.10	0.19	0.20	0.19	0.18	0.19	0.07
11Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
l1Mg	0.77	0.82	0.82	0.77	0.81	0.81	0.77	0.77	0.75	0.77	0.78	0.76
12Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
$12 \mathrm{Fe}^{2+}$	0.11	0.14	0.15	0.11	0.12	0.14	0.11	0.11	0.14	0.11	0.12	0.15
12Mn	0.02	0.02	0.01	0.02	0.02	0.01	0.02	0.02	0.02	0.02	0.02	0.01
12Ca	0.85	0.82	0.82	0.85	0.85	0.83	0.85	0.85	0.82	0.84	0.85	0.82
12Na	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02
12K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
/0	43.26	42.10	42.75	43.05	43.07	42.81	43.37	43.29	42.17	43.14	43.13	43.08
u	39.15	41.92	42.68	39.21	41.04	41.99	39.16	39.05	38.48	39.47	39.62	39.78
s	17.60	15.98	14.58	17.74	15.89	15.21	17.47	17.66	19.35	17.40	17.25	17.14
lg#	0.72	0.77	0.78	0.72	0.74	0.78	0.72	0.72	0.70	0.72	0.72	0.77

Sample	S35-6M2	S35-6M3	S35-7CR1	S35-7Cr1b	S35-7CR2	S35-7Cr2b	S35-7CR3	S35-7Cr3b	S35-7CR4	S35-7Cr4b
SiO,	52.21	51.64	52.65	52.65	52.95	52.86	52.24	52.18	52.70	52.15
TiO,	0.11	0.54	0.14	0.07	0.12	0.12	0.25	0.29	0.33	0.33
$^{2}$ Al <sub>3</sub> O <sub>3</sub>	0.55	2.64	0.98	0.58	0.85	0.75	1.43	1.91	1.60	1.75
FeO	11.19	9.39	10.47	10.71	10.07	9.84	9.11	8.91	8.74	8.52
$Cr_{2}O_{3}$	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.01	0.02	0.00
MnO	0.81	0.30	0.64	0.86	0.64	0.59	0.45	0.41	0.38	0.36
MgO	11.10	16.30	13.92	13.40	14.07	14.20	14.69	15.03	15.41	15.03
CaO	23.81	19.14	21.16	21.33	21.22	21.26	21.11	20.91	21.08	21.20
Na,O	0.10	0.15	0.26	0.23	0.25	0.27	0.29	0.29	0.28	0.31
<sup>z</sup> K,0	0.11	0.02	0.00	0.00	0.01	0.00	0.00	0.00	0.01	0.00
Total	66.66	100.12	100.22	99.84	100.18	99.89	99.57	99.94	100.55	99.65
TSi	1.98	1.91	1.97	1.98	1.98	1.98	1.95	1.94	1.94	1.94
TAI	0.02	0.09	0.03	0.02	0.02	0.03	0.05	0.06	0.06	0.06
$\mathrm{TF}\mathrm{e}^{3+}$	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MIAI	0.01	0.02	0.01	0.01	0.01	0.01	0.01	0.02	0.01	0.02
M1Ti	0.00	0.02	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01
$M1Fe^{3+}$	0.02	0.06	0.04	0.03	0.02	0.03	0.05	0.05	0.05	0.05
$M1Fe^{2+}$	0.34	0.01	0.18	0.21	0.18	0.17	0.12	0.09	0.08	0.09
MICr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
M1Mg	0.63	0.90	0.78	0.75	0.78	0.79	0.82	0.83	0.85	0.83
M2Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
$M2Fe^{2+}$	0.00	0.22	0.12	0.10	0.11	0.11	0.12	0.14	0.14	0.12
M2Mn	0.03	0.01	0.02	0.03	0.02	0.02	0.01	0.01	0.01	0.01
M2Ca	0.97	0.76	0.85	0.86	0.85	0.85	0.84	0.83	0.83	0.84
M2Na	0.01	0.01	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02
M2K	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Wo	48.96	38.76	43.00	43.52	43.16	43.24	43.07	42.59	42.46	43.22
En	31.76	45.92	39.36	38.04	39.82	40.19	41.70	42.59	43.19	42.64
Fs	19.28	15.32	17.64	18.44	17.02	16.57	15.23	14.82	14.35	14.14
$Mg^{*}$	0.65	0.79	0.73	0.71	0.73	0.74	0.77	0.79	0.79	0.79

ample	S28-31	S28-32	S28-33	S28-34	S28-41	S35-5-1	S35-52	S35-53	S35-54
0	56.08	51.81	50.89	51.78	50.30	43.46	42.90	42.95	42.76
$\mathbf{O}_2$	0.18	0.73	0.79	0.70	0.98	2.57	2.76	2.76	2.80
$1_2O_3$	1.16	3.74	4.59	3.66	5.01	10.82	11.22	11.39	11.02
90	8.50	10.92	11.00	10.23	11.77	12.59	12.28	12.60	12.62
$r_2O_3$	0.00	0.00	0.01	0.01	0.02	0.01	0.00	0.00	0.01
nO	0.40	0.33	0.30	0.34	0.29	0.24	0.27	0.25	0.27
gO	19.53	17.15	16.86	17.42	15.95	14.04	14.08	13.96	13.72
aO	11.65	11.38	11.55	11.54	11.64	11.19	11.31	11.26	11.23
$a_2O$	0.26	0.71	0.92	0.80	0.94	2.09	2.12	2.11	2.12
0	0.06	0.31	0.38	0.30	0.43	0.97	1.01	1.04	1.02
otal	97.82	97.08	97.29	96.78	97.33	97.98	97.95	98.32	97.57
Si	7.74	7.34	7.23	7.36	7.20	6.29	6.22	6.20	6.24
AI	0.19	0.62	0.77	0.61	0.80	1.71	1.78	1.80	1.76
$Fe^{3+}$	0.07	0.03	0.00	0.03	0.00	0.00	0.00	0.00	0.00
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
AI	0.00	0.00	0.00	0.00	0.05	0.14	0.13	0.14	0.13
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
$Fe^{3+}$	0.69	0.80	0.76	0.70	0.63	0.78	0.75	0.79	0.71
Ti	0.02	0.08	0.08	0.08	0.11	0.28	0.30	0.30	0.31
Mg	4.02	3.62	3.57	3.69	3.40	3.03	3.04	3.01	2.99
$Fe^{2+}$	0.22	0.46	0.55	0.49	0.78	0.75	0.74	0.74	0.83
Mn	0.05	0.04	0.04	0.04	0.04	0.03	0.03	0.03	0.03
Са	1.72	1.73	1.76	1.76	1.79	1.74	1.76	1.74	1.76
Na	0.07	0.20	0.24	0.22	0.22	0.26	0.24	0.26	0.24
Са	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Na	0.00	0.00	0.01	0.00	0.05	0.32	0.35	0.33	0.36
K	0.01	0.06	0.07	0.05	0.08	0.18	0.19	0.19	0.19
$\alpha/(M\sigma+Fe^{2+})$	0.95	0.89	0.87	0.88	0.81	0.80	0.80	0.80	0.78

Sample	H23	F12	F2	H3	F42	H7	S65	S15	F54	S10	S26	S28	S35
	dik A	dik A	dik A	dik A	dik A	dik A	dik A	dik B	dik B	dik B	dik B	dik B	dik B
SiO <sub>2</sub>	75.50	70.45	76.90	75.20	72.69	68.30	65.30	59.60	52.71	51.10	46.50	49.30	54.10
TiO <sub>2</sub>	0.13	0.29	0.08	0.06	0.24	0.32	0.39	0.51	1.05	0.66	0.91	0.96	0.72
$Al_2O_3$	12.45	14.1	12.32	11.9	13.33	14.85	15.15	15.2	16.58	16.85	16.5	15.95	17.60
Fe <sub>2</sub> O <sub>3</sub>	0.84	1.31	0.57	0.32	1.63	1.16	1.43	1.94	3.107	3.43	2.21	2.97	2.36
FeO	0.93	1.59	0.57	0.34	1.69	1.29	1.54	2.82	4.56	4.18	3.22	4.64	2.27
MnO	0.02	0.09	0.03	0.03	0.03	0.04	0.09	0.14	0.17	0.17	0.13	0.17	0.21
MgO	0.33	0.66	0.16	0.12	0.96	1.12	1.75	2.03	3.22	3.47	4.27	3.01	2.77
CaO	1.02	3.96	0.63	1.82	0.56	1.24	2.05	3.90	8.05	4.48	8.43	9.09	6.15
Na <sub>2</sub> O	3.06	1.38	3.53	0.17	0.66	3.98	4.68	2.37	2.77	4.76	6.13	2.82	1.00
$K_2O$	4.78	5.87	5.14	8.25	7.93	4.22	3.96	3.93	3.82	3.61	0.89	3.45	9.35
$P_2O_5$	0.02	0.10	0.01	0.01	0.08	0.11	0.13	0.17	0.40	0.22	0.30	0.38	0.25
LOI	0.80	0.08	100.12	1./3	0.15	2.40	3.20	08.65	3.40	0.07	9.78	5.58	2.98
Total	99.88	207.4	117.5	99.95	99.95	99.03	99.07	98.05	99.84 502	529	99.20	98.33	99.70
Da Ph	40/	327.4 140	165	200	173	127	413	112	124	920 83	22	116	208
NU Sr	103 50	54.10	105	209	34.60	77.80	90 83.00	07.60	124	125.00	160 50	527.00	252 462.00
Ga	13.10	*	*1.40	10.50	13.40	15.10	14.80	15 20	17 70	125.00	16.80	18.90	16.00
Nh	14.80	10.83	17.63	12.10	14 50	15.10	12.60	13.20	23.60	10.00	20.3	24 50	13.00
Hf	3.60	*	*	2.30	3.50	4.10	4.40	4.30	4.90	3.30	3.50	5.60	3.40
Zr	107	129	101	62	129	148	163	155	191	124	149	223	123
Y	13.70	20.50	22.60	13.90	13.70	15.50	16.10	18.10	22.90	22.40	29.10	23.50	20.70
Th	26.9	14.41	47.60	14.55	15.45	14.40	14.40	13.30	14.90	9.36	5.65	16.00	10.80
U	5.36	*	*	3.09	2.95	4.32	3.59	3.37	3.94	2.65	2.22	4.55	3.35
Cr	10	0	0	10	10	10	10	10	230	20	370	160	10
Ni	7	2	2	6	5	5	7	6	63	10	109	50	11
Co	4.60	*	*	0.50	3.90	4.80	4.30	9.60	23.80	16.20	11.10	22.10	9.40
V	19	*	*	5	22	43	62	96	196	167	168	226	206
Zn	23	*	*	7	20	21	30	27	91	375	46	44	121
Cd	*	*	*	*	*	*	30	27	*	375	46	44	121
W	3.00	*	*	1.00	16.00	2.00	4.00	4.00	5.00	3.00	12.00	2.00	3.00
Та	1.80	*	*	1.50	1.40	1.30	1.10	1.20	1.30	0.70	1.50	1.60	0.90
Se	*	*	*	*	*	*	1.10	1.20	*	0.70	1.50	1.60	0.90
Cs	3.81	*	*	2.40	1.19	3.01	1.19	5.78	5.76	0.50	0.23	5.12	1.34
La	33.60	*	*	30.20	28.80	34.10	35.10	32.70	35.50	23.40	20.30	35.70	27.00
Ce	58.20	40.50	70.40	52.70	53.00	60.30	63.20	58.70	66.90	46.10	39.50	67.70	52.40
Pr	5.66	*	*	5.58	5.67	6.18	6.75	6.43	7.73	5.44	4.81	7.80	6.25
Nd	17.60	*	*	18.40	18.20	20.30	23.70	22.20	28.10	21.20	19.10	29.70	23.80
Sm E	2.92	* *	~ *	3.48	3.13	3.51	4.03	4.21	5.//	4.31	4.16	5.86	4.92
Eu	0.50	*	*	0.74	0.51	0.85	0.79	1.04	1.47	1.15	1.19	1.45	1.28
Са ТЪ	2.07	*	*	2.90	0.38	5.54 0.45	5.8 0.47	4.15	5.45 0.70	4.54	4.09	0.81	4.65
10 Dv	0.57	*	*	0.41	0.58	0.45	0.47	2 20	1.29	0.70	0.78	4.20	2.06
Но	0.46	*	*	0.50	0.43	0.53	0.58	0.65	4.38	4.00	1.02	0.89	0.78
Fr	1.47	*	*	1.66	1.56	1.71	2.00	2.02	2.44	2.45	3.40	2.50	2 31
Tm	0.25	*	*	0.26	0.23	0.26	0.33	0.29	0.34	0.37	0.49	0.35	0.34
Yh	1.73	*	*	1.85	1.55	1.87	2 20	2.01	2.13	2 40	3.02	2 24	2 14
Lu	0.28	*	*	0.30	0.25	0.32	0.38	0.33	0.35	0.36	0.45	0.36	0.33
A/CNK	1.03	0.89	0.98	0.95	1.25	1.11	0.96	1.00	0.58	0.85	0.63	0.64	0.77
(La/Yb)n	13.19	*	*	11.09	12.62	12.39	10.84	11.05	11.32	6.62	4.57	10.83	8.57
(Eu/Eu*)	0.54	*	*	0.69	0.51	0.75	0.61	0.75	0.79	0.77	0.82	0.74	0.79
Ce/Yb	33.64	*	*	28.49	34.19	32.25	28.73	29.20	31.41	19.21	13.08	30.22	24.49
Ta/Yb	1.04	*	*	0.81	0.90	0.70	0.50	0.60	0.61	0.29	0.50	0.71	0.42
Zr/Sm	36.64	*	*	17.82	41.21	42.17	40.45	36.82	*	*	*	*	*
Nb/Ta	8.22	*	*	8.07	10.36	12.00	11.45	11.50	*	*	*	*	*
Ba/Rb	*	*	*	*	*	*	*	*	4.78	6.37	7.28	4.81	10.63
Rb/Sr	*	*	*	*	*	*	*	*	0.25	0.66	0.13	0.22	0.55

Tab. 6 The whole-rock geochemical data from acidic and intermediate-basic dikes (major elements in wt. %, trace elements in ppm)

dik A: acidic dikes, dik B: basic-intermediate dikes

\* - not determined



The clinopyroxene of the basic dikes corresponds to diopside–augite (Fig. 6a). The Mg<sup>#</sup> of the clinopyroxene phenocrysts vary a little around 0.7, and the average value correlates well with the whole-rock compositions. Some phenocrysts show oscillatory zoning in Al, Ti and Mg contents, consistent with the crystal growth and/or the element diffusion in liquids with different compositions and temperatures (Ridolfi et al. 2008). According to the classification of Le Bas (1962), the pyroxenes show subalkaline affinity (Fig. 6b). The association of diopside–augite clinopyroxene is characteristic of shoshonitic rocks as well (Morrison 1980).

According to classification scheme of Leake et al. (1997) and Hawthorne and Oberti (2007), amphiboles of intermediate dikes are mainly magnesiohornblendes, whereby those from basic dikes correspond to magnesiohastingsite ( $^{VI}Al < Fe^{3+}$ ; Fig. 7a). Tremolite probably crystallized as a subsolidus phase since it falls outside the limit of igneous amphibole.

Based on earlier studies (Chappell and White 1974; White and Chappell 1983; Wyborn et al. 1981; Clemens and Wall 1984), the presence of calcic amphiboles in granitoid rocks indicates I-type affinity. Fig. 7b shows that the compositions of amphibole from intermediate dikes are consistent with crystallization from subalkaline magmas, while those from the basic dikes have higher TiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> contents and thus fall into the transitional subalkaline–alkaline domain (Molina et al. 2009). The compositional differences between amphiboles from both rock types may imply a difference in the physicochemical conditions of magmas from which they crystallized.

#### 6. Geochemistry

Geochemical compositions of the studied dikes are listed in Tab. 6. Using the  $Zr/TiO_2$  vs. silica diagram (Winchester and Floyd 1977), the silicic dikes can be classified as rhyolite, rhyodacite and dacite, while the intermediate–basic rocks show phonolite, basanite and trachyandesite composition (Fig. 8).

In the A/CNK [molar  $Al_2O_3/(CaO+Na_2O+K_2O)$ ] versus Zr + Nb + Ce + Y diagram (Condie et al.1999; Fig. 9), the acidic dikes plot as metaluminous to slightly peraluminous. The intermediate-basic dikes fall in the metaluminous field, some far from the discrimination line of the metaluminous and peraluminous domains. This diagram shows that the studied dikes are of I- rather than A-type nature.

**Fig. 5** Biotite compositions from intermediate dikes. **a** – classification diagram of Rieder et al. (1998). **b** –  $Al_2O_3$ –FeO<sub>T</sub> diagram (Abdel-Rahman 1994) discriminating the likely geotectonic setting of biotites. **c** – Fe<sup>2+</sup>/(Fe<sup>2+</sup>+Mg) vs. Al<sup>IV</sup> diagram (Jiang et al. 2002).



**Fig. 6** Composition of clinopyroxenes from the basic dikes.  $\mathbf{a}$  – Wo–En–Fs classification diagram (Morimoto 1988).  $\mathbf{b}$  – Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> diagram (Le Bas 1962) discriminating the nature of the pyroxene host rocks.



Fig. 7a-b Classification diagram for amphiboles (Leake et al. 1997; Hawthorne and Oberti 2007). c – Composition of studied amphiboles in the Al<sub>2</sub>O<sub>3</sub> vs. TiO<sub>2</sub> binary diagram (Molina et al. 2009).



Fig. 8  $Zr/TiO_2$  vs. SiO<sub>2</sub> diagram for nomenclature of the studied dikes (Winchester and Floyd 1977).



**Fig. 9** A/CNK  $[Al_2O_3/(CaO+Na_2O+K_2O) molecular ratio]$  vs. Zr + Nb + Ce + Y (Condie et al. 1999), showing the nature of the studied dikes.

100 100 (a) High-K calc-alkaline (b) and shoshonitic Shoshonitic 88 10 X X Ce/Yb 10 Th 1 Calc-Alkaline , Calc-Alkaline 0.1 Tholeiitic Dacite-(Basaltic) Basalt Rhyolite Andesite/ 0.01 60 40 20 0 0.0 Ta/Yb Co

Fig. 10 Co vs. Th (Hastie et al. 2007) (a) and Ce/Yb vs. Ta/Yb diagram (Pearce 1982) (b) which document the shoshonitic affinity of the studied dikes.

The Th–Co diagram (Hastie et al. 2007; Fig. 10a) has been designed to replace a classic  $SiO_2$ –K<sub>2</sub>O plot of Peccerillo and Taylor (1976) for altered igneous rocks. The studied dikes plot in the high-K calc-alkaline and shoshonitic field. Likewise, in Ce/Yb versus Ta/Yb diagram (Pearce 1982; Fig. 10b), most of the samples fall in the shoshonitic domain.

All indicators such as (1) high alkali content ( $K_2O + Na_2O > 5$  wt. %: Turner et al. 1996), (2) LILE enrichment relative to HFSE, and (3) low TiO<sub>2</sub> content (< 1.2 wt. %: Morrison 1980) document that the Kuh-e Dom dikes have shoshonitic affinity. Furthermore, Zr/Hf ratios range between 27 and 42, consistent with similar ratios in shoshonitic igneous rocks (Carmichael et al. 1996; Müller and Forrestal 1998).

In chondrite-normalized REE plots (Sun and Mc-Donough 1989; Fig. 11a), the acidic dikes display lower HREE contents and higher degree of LREE/HREE enrichment than the intermediate–basic dikes do. The silicic dikes (Fig. 11b) show nearly parallel trends, with moderate LREE enrichment ((La/Yb)<sub>n</sub> = 11–13) and sizeable negative Eu anomalies (Eu/Eu\* = 0.51–0.75). Also the REE patterns for intermediate–basic dikes are subparallel, but the negative Eu anomalies are shallower (Eu/Eu\* = 0.79–0.90) and the patterns less steep ((La/Yb)<sub>n</sub> = 5–11).

Primitive mantle-normalized (Sun and McDonough 1989) patterns for silicic dikes (Fig. 12a) are largely parallel and show LILE enrichment relative to HFSE, as well as negative anomalies of Nb (Nb/Nb\* ~0.12), Ta (Ta/Ta\* ~0.3), Sr (Sr/Sr\* ~0.13), P (P/P\* ~0.19) and Ti (Ti/Ti\* ~0.14). Similar features show the intermediate–basic dikes (Fig. 12b) even though the negative anomalies are less conspicuous (Nb/Nb\* ~0.25, Ta/Ta\* ~0.3 and Ti/Ti\* ~0.48). The Sr contents of these dikes (except two samples which are altered to spilite) do not show any depletion.

#### 7. Discussion

#### 7.1. Dikes origin

7.1.1. Silicic dikes

We discuss three models for the generation of silicic dikes: (1) fractional crystallization of basic magma parental to the intermediate–basic dikes with, or without, assimilation, (2) origin from residual melt during the final stages of crystallization of the granite host and (3) partial melting of lower crust. Were the silicic dikes generated by fractionation of more basic melts, the REE patterns of all rock types should be parallel but they are not (Fig. 11a). Moreover, the relatively large volume of silicic dikes compared to intermediate-basic dikes argues against such a model.

Silicic dikes have HREE contents significantly lower, but their LREE contents fall within, the span observed in their igneous host rocks (shaded area in Fig. 11b). Assuming that the silicic dikes are products of evolved residual granitic melt, they should show higher LREE content relative to the parental magma (Rollinson 1993) due to the remarkably greater abundance of amphibole and biotite in the host granite, but this is not the case (Fig. 11b).

In primitive mantle-normalized spider plots, the silicic dikes show negative anomalies in Ti, P, Sr, Nb and Ta. These are important features of fractionated granitic magmas, e.g. in the Lachlan Orogenic Belt (White and Chappell 1983; Chappell and White 1992). Experimental studies show that silica-rich (>56 wt. % SiO<sub>2</sub>) and high-K, I type granitoid magmas can be derived from partial melting of older metaigneous rocks in the continental crust (Roberts and Clemens, 1993; Patiño Douce and Mc-Carthy 1998; Yardley and Valley 1997). Because of their low K<sub>2</sub>O contents, metabasaltic rocks of all kinds are unsuitable as sources unlike the hydrous, calc-alkaline to high-K calc-alkaline, mafic to intermediate metamorphic rocks (Roberts and Clemens 1993).

Melting of pelitic crust alone could not produce metaluminous magma. There is no evidence on the extensive magmatic fractionation among the samples. In the  $Al_2O_3$ + FeOt + MgO + TiO<sub>2</sub> vs.  $Al_2O_3/(FeOt + MgO + TiO_2)$  diagram (Fig. 13; Patiño Douce 1999), most of the samples plot in the field of greywacke-derived melts. Therefore it is seen that partial melting of a quartzo-feldspathic crust was probably the main mechanism for formation of silicic dikes.

#### 7.1.2. Intermediate-basic dikes

The more basic dikes show shoshonitic affinity with LILE and LREE enrichment. Some authors interpret subcontinental lithospheric mantle, or asthenospheric mantle enriched in incompatible elements by earlier subduction, as the principal sources of shoshonitic and associated high-K calc-alkaline magmas (Aldanmaz et al. 2000; Seghedi et al. 2004). Others have emphasized magma fractionation and crustal assimilation as the dominant processes (Meen 1987; Feeley and Cosca 2003).

Experimental data indicate that the K-rich magma could not form by low-degree (2–2.5%) partial melting of lherzolite at less than 2–3 GPa (70–100 km) as it can



**Fig. 11** Chondrite-normalized rare earth elements patterns of Kuh-e Dom dikes (Sun and McDonough 1989) (**a**) and the comparison of acidic (**b**) and intermediate-basic dikes (**c**) with the host-rock patterns (gray field).

only yield basanitic magma with  $SiO_2$  contents inconsistent with the criteria for shoshonites (Wang et al. 1991). Crustal contamination alone also cannot generate K-rich magma from partial melt originally derived from pyrolite (Wang et al. 1991).





Fig. 12 Primary mantle-normalized (Sun and McDonough 1989) trace-element patterns of Kuh-e Dom dikes, acid (a) and intermediate-basic (b).

Many researchers proposed that the K-rich metasomatized lithospheric mantle is the main source of shoshonitic and mafic K-rich rocks in continental arcs (e.g. Meen and Eggler 1987; Lange and Carmichael 1991; Edwards et al. 1991; Wallace and Carmichael 1992). The enrichment of the mantle wedge in incompatible elements can be related to (Rock 1991): (1) involvement of an anomalous, metasomatized potassium-rich lithospheric mantle domain, or (2) recycling of a crustal component in subduction zones (e.g. Tatsumi and Eggins 1995; Jahn et al. 1999). Moreover, the presence of water-bearing minerals such as biotite and hornblende in the lithospheric mantle indicates that it had been metasomatized (Hawkesworth et al. 1990) as low Nb/U ratios (3-17) do (Griffin et al. 1988; O'Reilly and Griffin 1988; Maury et al. 1992). This is consistent with the previous experimental results (Edgar et al. 1976; Mengel and Green 1989). The experiments of Wyllie and Sekine (1982) demonstrated that reaction between hydrous fluids/melts and mantle peridotite produce hybrid mantle which could subsequently yield potassic melts (Jiang et al. 2005).

An important feature of subduction-related mantle metasomatism is the formation of hydrous mineral phase(s) phlogopite and/or amphibole (Beccaluva et al. 2004). Müller and Groves (1997) suggested that the likely mineral in the source is phlogopite for many potassic rocks, and that melting of a phlogopitebearing peridotite may produce potassium-rich parent magma. Zhang et al. (2008) argued that the presence of phlogopite or pargasitic hornblende in the source can lead to formation of K-rich magmas. Indeed, the experiments confirmed the stability of these potassium- and water-bearing minerals at < 30kbar and T < 1050 °C (Edgar et al. 1976; Mengel and Green 1989).

Given the differences in

compatibility relative to phlogopite and amphibole, the ratios of some incompatible elements can be used to constrain which hydrous phase existed in the melted lithospheric source (Furman and Graham 1999; Yang et al. 2004; Xie et al. 2006). For example, melts in equilibrium with phlogopite should show higher Rb/Sr (> 0.1) and lower Ba/Rb (< 20) ratios than those from amphibole-bearing sources at the same degree of partial melting (Furman and Graham 1999). As Ba is partitioned into phlogopite to a greater extent than Rb (Ionov et al. 1997), partial melts, which leave phlogopite in the residue, will acquire a low Ba/Rb ratio. On the other hand, significant removal of phlogopite from the residue by (higher degree) partial melting of the veins will tend to increase the Ba/Rb ratios (Kurt et al. 2008).

Figure 14 shows that high-K samples from the Kuh-e Dom intermediate–basic dikes tend to display low Ba/ Rb and variably high Rb/Sr ratios, which support a low-degree partial melting of phlogopite-bearing veined lithospheric mantle source. Moreover, the studied dike samples have constantly high Yb concentrations exceeding 1.8 ppm and thus do not support amphibole as residual mineral phase in the mantle source (Guo et al. 2005). However, the degree of enrichment of middle REE relative to heavy REE such as Yb depends mainly on garnet existence as a residual phase during melting because HREE are preferentially retained by this mineral (Zhang et al. 2008). Hence, minor changes in Sm/Yb ratio (1.3–2.7) in the studied intermediate–basic dikes underline lack of garnet in the mantle source.

Although phlogopite is not a typical mantle mineral and its presence in the upper mantle reveals compositional anomalies (Peccerillo 2003), it seems that a phlogopite-bearing lithospheric mantle metasomatized by subduction-related fluids or melts could be a potential melt source in our case (Conticelli and Peccerillo 1992). This small-scale heterogeneity is commonly considered to be due to the presence of a vein network permeating the upper-mantle peridotite (Foley 1992; Perini et al. 2004; Krmíček et al. 2011).

The veins are inferred to be rich in clinopyroxene together with accessory phases, such as phlogopite, titanite, and oxides, uncommon in normal peridotitic mantle. The low degrees of partial melting of such a veined phlogopite-bearing mantle would allow almost pure vein material to melt, generating highly potassic magmas. At higher degrees of melting, the same source would allow increasing amounts of lherzolite wallrock to enter the melt, generating magmas with lower enrichments in incompatible elements.

Potassium-bearing clinopyroxenes are known to exist in the deep mantle at P = 5-7 GPa and T > 1200 °C (Tsuruta and Takahashi 1998). However, melting of deep mantle lithologies in the presence of residual K-clinopyroxene appears to be unlikely in case of the Kuh-e Dom dikes because (1) it would result in near horizontal trends of CaO versus MgO and SiO<sub>2</sub> in the melt which are not observed, and (2) the HP–HT mantle melting conditions required for the generation of alkaline magma (1300–1500 °C; Tsuruta and Takahashi 1998) would result in significantly larger volumes of melts which do not appear to exist in the studied region.

The range of variation in the trace element (LREE/ HFSE) characteristics of the magmas thus might be attributed to small-scale mantle heterogeneities, variable role for accessory phases in the residue (e.g., phlogopite and apatite) or the difference in the relative proportion



**Fig. 13**  $Al_2O_3$ +FeOt+MgO+TiO<sub>2</sub> vs.  $Al_2O_3$ /(FeOt+MgO+TiO<sub>2</sub>) (Patiño Douce 1999) diagram indicating the partial melting of a quartzo-feld-spathic crustal source.



Fig. 14 Low Ba/Rb and high Rb/Sr ratios, showing phlogopite-bearing mantle source of intermediate–basic dikes (Furman and Graham 1999).

of veins and peridotitic wall-rock components, which contributed to the melt (i.e., degree of partial melting). The variable K-contents probably reflect melting of different amounts of phlogopite, or variable degree of partial melting of a phlogopite-bearing melt source.

#### 7.2. Tectonic setting

Potassium-rich rocks are found in several tectonic environments, such as cratons, post-collisional setting, active orogenic belts and, to a lesser extent, oceanic intraplate



Fig. 15 Distribution of the selected samples in the geotectonic discrimination diagrams of Condie (1989) (a) and Müller and Groves (1997) (b–c).

settings (Müller and Groves 1997). Shoshonitic rocks are generally associated with many continental and oceanic arcs (Peccerillo 1992).

The studied dikes show enrichment in LILE and LREE, relative to HREE and HFSE, with negative anom-

alies in Ta, Ti and Nb. The average Ba/Nb and Rb/Nb ratios are 37.5 and 10.5 in acidic dikes and 62.5 and 7.6 in the intermediate-basic dikes, respectively. Such high values are typical of potassic igneous rocks formed at active continental margins (Müller and Groves 1997) because HFSE tend to remain in the subducted plate, whereas the LILE are easily transferred into the mantle wedge (Pearce 1983; Tatsumi and Eggins 1995). By comparison, these ratios are relatively low in the OIB (~7.3 and ~0.6) and MORB (~2.7 and ~0.24; Sun and McDonough 1989; Harangi 1994; Carmichael et al. 1996). Moreover, it is noted that the all Kuh-e Dom shoshonitic dikes plot within the fields of continental arcs in the La/Yb-Th/ Yb tectonomagmatic discrimination diagram (Condie 1989; Fig 15a). The low Hf/Sm ratio (~1) in the studied dikes also indicates a subduction-related orogenic setting (Lafleche et al. 1991). Finally, the Zr/Hf ratios (27–37: acidic dikes, 33.6-42.6: intermediate-basic dikes) are consistent with potassic igneous rocks in continental-arc settings (Carmichael et al. 1996).

Müller and Groves (1997) suggested various discrimination diagrams for distinguishing between different tectonic settings of potassic rocks, especially active continental margins from post-collisional settings. All of Kuh-e Dom shoshonitic dike compositions classify as arc-related (Fig. 15b); in Fig 15c the silicic dikes fall in the continental-arc field while the intermediate–basic dikes seem post-collisional.

In summary, the combination of field, petrographic and geochemical data of the Kuh-e Dom shoshonitic dikes demonstrates that they formed at late stage of development of an active continental margin. This is in accord with previous reports on the petrogenesis of igneous rocks in the Urumieh-Dokhtar magmatic belt (Berberian and King 1981; Mohajjel et al. 2003; Shahabpour 2007). Indeed, recent tectono-magmatic syntheses have suggested that the Eocene magmatic event may have been a consequence of geometrically complex subduction of Neotethys oceanic crust underneath Central Iran. The acidic dikes which intruded the Kuh-e Dom pluton were emplaced in a transitional subduction to syn-collisional setting whereas the intermediate-basic dikes have developed after a short time gap, probably in the postcollisional setting (Technoexport 1981).

Recently, geologists have paid more attention to the dike swarms produced after the large-scale orogenic granite magmatism (Dini et al. 2002; Jiang et al. 2006), which have been termed the post-batholith dike swarms (Allen 2000) or post-orogenic complexes (Luo et al. 2006). It is proposed that such dike swarms are produced during the switch of the regional tectonic stress field from compression to extension. Therefore they could be marking the end of the orogeny similar to Taihangshan, Yanshan, Eastern Kunlun and Tianshan orogenic belts (Luo et al 2008).

# 8. Conclusions

Based on petrographic and geochemical data from the Eocene Kuh-e Dom shoshonitic dikes, the following conclusions were reached:

- 1. All shoshonitic dikes can be subdivided into two groups: silicic (porphyric microgranodiorites, porphyric microgranites and aplites) and intermediate–basic (chemically ranging from phonolite to trachyandesite and basanite).
- 2. The silicic dikes are subaluminous and show shoshonitic affinity, with Ti, P, Sr and Nb depletion in Primitive-mantle normalized spiderplots indicating generation by partial melting of a quartzo-feldspathic source.
- 3. The intermediate-basic dikes have a metaluminous and shoshonitic nature. The high Rb/Sr with low Ba/Rb ratios document that their parental magma originated by partial melting of a phlogopite-bearing enriched mantle.
- 4. The mineral chemistry and whole-rock enrichments in LILE and LREE over HFSE and HREE, low Hf/ Sm, as well as high Ba/Nb, Rb/Nb and Zr/Hf ratios indicate that Kuh-e Dom dikes formed during subduction of oceanic lithosphere beneath central Iran continent. The silicic dikes intruded during the late subduction and the intermediate-basic dikes soon thereafter.

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