# Molybdenite-tungstenite association in the tungsten-bearing topaz greisen at Vítkov (Krkonoše-Jizera Crystalline Complex, Bohemian Massif): indication of changes in physico-chemical conditions in mineralizing system

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At Vítkov, sparse molybdenite occurs within tungsten mineralization hosted by topaz greisen in orthogneiss in the envelope of the Variscan Krkonoše–Jizera granite Pluton (northern Bohemian Massif). Mineralogical study showed that sulfide mineralization started with precipitation of arsenopyrite followed by molybdenite, tungstenite, transitional Mo- and W-dominated disulfides and concluded by pyrite. Textural relationships between molybdenite and tungstenite imply that tungstenite was formed during several stages related to molybdenite bending and fracturing. Laser Ablation Inductively Coupled Mass Spectrometry (LA-ICP-MS) analyses of Re-poor (< 0.3 ppm) molybdenite showed extreme concentrations of W (up to 26 558 ppm) accompanied by Ag, As, Bi, Pb, Se, Te and other metals. Electron microprobe analyses of inclusions-free molybdenite confirmed the abundance of W (~0.5 wt. %) and tungstenite showed ~4 wt. % Mo, indicating a substitution of Mo<sup>4+</sup> for W<sup>4+</sup>. Stability and phase relationships between molybdenite and tungstenite and locally identified transitional Mo- and W-dominated disulfidic phases suggest that tungstenite crystallization was triggered by a decrease in  $fO_2$  below WO<sub>2</sub>–WO<sub>3</sub> buffer that followed after molybdenite precipitation. Tungstenite zoning and sharp tungstenite–molybdenite contacts indicate disequilibrium during their formation.

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# 1. Introduction

It was documented by many authors (e.g. Cook et al. 2009, 2013; Large et al. 2009, 2014; Sung et al. 2009; Lin et al. 2011; Thomas et al. 2011; Pašava et al. 2013, 2015, 2016) that Electron Microprobe Analysis (EMPA) and Laser Ablation Inductively Coupled Mass Spectroscopy (LA-ICP-MS) spot analyses and mapping are very useful for documentation and interpretation of trends of trace-element distribution in sulfides such as pyrite, arsenopyrite and sphalerite. Blevin and Jackson (1998) used LA-ICP-MS to measure a number of elements in molybdenite containing both Bi-mineral inclusions and Au. They evidenced presence of Pb, Bi, Te and other elements and suggested for the first time that variation in their abundance could potentially be a valuable petrogenetic tool. This outcome was supported by Norman et al. (2004) who highlighted the importance of molybdenite trace-element data, notably Bi, for fingerprinting mineralization in granitic settings. Recently, Ciobanu et al. (2013), confirmed the usefulness of this kind of study for identification of discrete events in young magmatic-hydrothermal systems on example of two Au and Au-Cu-(Mo) porphyry-style deposits.

As an important carrier of Re, molybdenite has been proven to be a useful mineral in Re–Os geochronology (Stein et al. 2001). The method yields an age of mineralization and it is robust with respect to subsequent overprinting (e.g. Stein et al. 2003; Selby and Creaser 2004) except when subject to supergene alteration (Mc-Candless et al. 1993). However, Košler et al. (2003) have documented heterogeneity within single grains of molybdenite caused most likely by decoupling of Re and Os which could, in turn, affect the accuracy of Re–Os ages. They suggested the need for homogenization of substantial quantities of molybdenite to capture the bulk Os/Re of the sample.

The Bohemian Massif is known for numerous Variscan Au, Sn–W, U and base-metal deposits located in different geotectonic units and settings. These mineralizations are sometimes accompanied by molybdenite, which is much more widespread in granite-related ore deposits (Drábek et al. 1993). No tungstenite was ever reported from this type of deposits in the Bohemian Massif. First study on chemistry of Re and Se in molybdenite in the Bohemian Massif was carried out by Kvaček and Trdlička (1970) and followed by Kvaček et al. (1983), and Drábek et al. (1989 and 1993). Košler et al. (2003) reported inhomogeneously distributed Re in molybdenite from several gold deposits in the Bohemian Massif.

In this paper we present new mineralogical, LA-ICP-MS and EMPA data for Re-poor molybdenite and associated transitional W–Mo/Mo–W disulfides from Vítkov (Krkonoše–Jizera Crystalline Complex, Bohemian Massif). We demonstrate how different physico-chemical conditions during mineralizing processes are reflected by

chemical composition of molybdenite and related sulfides and also by the presence of impurities in molybdenite.

## 2. Geological setting and mineralization

Topaz-bearing greisens belong to exotic rocks in the West Sudetes domain which is mostly considered a part of the Saxothuringian Zone (e.g. Kosmat 1927; Behr et al. 1984; Franke et al. 1993; Mikulski and Stein 2011; Mochnacka et al. 2015) of the Bohemian Massif located at the eastern termination of the European Variscan Belt (Fig. 1a). The West Sudetes in the northern margin of the Bohemian Massif (Fig. 1a) are composed of a heterogeneous mo-



Fig. 1a – Position of the Bohemian Massif and major tectonostratigraphic zones of the Variscan orogenic belt in Europe.  $\mathbf{b}$  – Simplified geological map of the Krkonoše–Jizera Pluton within the West Sudetes crystalline complexes (adapted from Chrt et al. 1968); associated mineral deposits and occurrences are also shown.  $\mathbf{c}$  – Detailed geological map of the Vítkov and Růžek quartz–topaz greisen localities.

saic of pre-Permian basement fragments comprising the Krkonoše-Jizera orthogneisses representing deformed and metamorphosed granitoids of the Lusatian Batholith of Late Proterozoic up to Ordovician age (Kozlowska 1956; Karwowski 1973) and Neoproterozoic to Lower Paleozoic low-grade metamorphic complex dominated by phyllites, and metavolcanic rocks (Kachlík and Patočka 1998). In the middle of these heterogeneous rock assemblages is situated the Krkonoše-Jizera granite Pluton (Fig. 1b) of Variscan age (Žák et al. 2013). Before the Krkonoše-Jizera granite Pluton emplacement, the Variscan tectonic history of these rocks involved oceanic subduction and blueschist-facies metamorphism dated at around 360 Ma, followed by collision, nappe stacking and widespread greenschist-facies metamorphism at around 340 Ma (Maluski and Patočka 1997; Farvad and Kachlík 2010).

Biotite and two-mica orthogneisses are strongly foliated rocks that were probably the parental material of the locally intensely metasomatized rocks containing quartz and topaz with white mica, tourmaline, traces of sulfides (molybdenite and arsenopyrite), wolframite and cassiterite (Karwowski 1973 and 1977). Similar mineral association is known from the interior of the Krkonoše– Jizera granite Pluton (Fig. 1b).

Stratabound quartz-topaz greisen bodies are located on both sites of the Polish–Czech border (Fig. 1b). In Poland, quartz-topaz greisens occur between leucogranites and mica schists between Rębiszów and Gierałtówek settlements (Mlądz–Kamień–Pobiedna Zone, Karwowski 1973, 1977). It is a narrow E–W trending zone, about a hundred meters wide and more than 10 km long, parallel to well-known strata-bound Nové Město p. Smrkem–Gierczyn–Krobica Variscan tin mineralization (Fig. 1b).

On the Czech territory, similar greisens were found near Růžek and Vítkov settlements (Watzernauer 1940; Klomínský et al. 2003; Vondrovic et al. 2015; Fig. 1c). On the both sides of the Polish–Czech border, the only remnants of the former outcrops are loose greisen blocks that occur parallel to other geological structures (metamorphic foliation) of the Krkonoše–Jizera orthogneisses, 3–10 km away from the Krkonoše–Jizera Pluton's northern margin (Fig. 1b). At Vítkov, dominant wolframite (ferberite) accompanied by accessoric scheelite, pyrite, molybdenite, arsenopyrite, native bismuth, rutile (with increased Nb and Ta) and zircon was reported by Sidorinová and Dobeš (in print).

## 3. Methods

We prepared polished sections from selected mineralized samples containing rather rare molybdenite and then used field-emission scanning electron microscope Tescan Mira3 GMU fitted with energy-dispersive spectrometer Oxford Instruments X-Max 80 (80 mm<sup>2</sup> active area) at the Czech Geological Survey for more detailed mineralogical study. Quant optimization was performed on the cobalt standard. Synthetic factory standards were used (FeS<sub>2</sub> for S, metallic Mo and metallic W) for qualitative analysis. Analytical conditions during the analysis were ~ 7 nA absorbed current at an accelerating voltage of 20 kV. Analytical results were processed in the AZtec software version 3.0 (Oxford Instruments).

The molybdenite trace-element data were obtained using an Element 2 high-resolution ICP-MS coupled with a 213-nm NdYAG UP-213 laser-ablation system at the Institute of Geology of the Czech Academy of Sciences, v.v.i. The following isotopes were measured at the two mass-resolution modes in order to avoid isobaric and polyatomic interferences: (1) the low mass-resolution mode (m/ $\Delta$ m = 300) was used for <sup>57</sup>Fe, <sup>75</sup>As, <sup>77</sup>Se, <sup>90</sup>Zr, <sup>93</sup>Nb, <sup>107</sup>Ag, <sup>118</sup>Sn, <sup>121</sup>Sb, <sup>125</sup>Te, <sup>182</sup>W, <sup>187</sup>Re, <sup>197</sup>Au, <sup>208</sup>Pb and <sup>209</sup>Bi, while (2) medium-resolution (m/ $\Delta$ m = 4 000) mode was required for <sup>56</sup>Fe, <sup>59</sup>Co, <sup>60</sup>Ni, <sup>63</sup>Cu and <sup>66</sup>Zn. More detailed description of the method was given in Pašava et al. (2016).

Electron-microprobe analyses (EMPA) were performed on the Cameca SX-100 instrument at the Joint Laboratory of Electron Microscopy and Microanalysis of the Masaryk University and the Czech Geological Survey (Brno, Czech Republic). The measurements were carried out using a wave-dispersion mode under the following conditions: accelerating voltage of 25 kV, beam current of 20 nA and the beam diameter of 1 µm. Natural and synthetic standards were employed (FeS<sub>2</sub> for Fe and S, pararammelsbergite for Ni and As, with metallic Co, Mo, W and Re). Used lines were: K<sub>a</sub> for Co, Fe, Ni and S,  $L_{a}$  for Re and W, and  $L_{B}$  for As and Mo. For molybdenite, peak count times were 20 s for Mo, W and S, and 120 s for Re. Background count times were 10 s for Mo and W, 20 s for S and 60 s for Re. For arsenopyrite, count times were 20 s on peak (Fe, Co, Ni, As and S) and 10 s on background.

#### 4. Results

#### 4.1. Mineralogy of major sulfides

Molybdenite in analyzed sample from greisen-type mineralization at Vítkov occurs mostly as fine- to mediumgrained, lamellar aggregates or fragments of lamellae with common kinks and bending (Fig. 2a). Molybdenite is younger than arsenopyrite (Fig. 2a). Similarly, tungstenite is younger than arsenopyrite and molybdenite. It forms nano- to micron-sized inter-lamellar fillings in



200 µm

by fragmented arsenopyrite (Aspy).

Sample grain size		Co	Ni	Cu	Zn	As	Se	Zr	Nb	Ag	Sn	Sb	Te	W	Re	Au	Pb	Bi
	*IDL 55 μm	0.2	22	0.5	0.5	0.5	1.:	5 0.07	0.10	0.04	0.3	0.05	0.4	0.10	0.03	0.01	0.05	5 0.03
Vítkov	fibrous fragment	42	<idl< td=""><td>33</td><td>7.2</td><td>1774</td><td>211</td><td>3.4</td><td>6.5</td><td>37</td><td>5.3</td><td>6.8</td><td>3.8</td><td>26558</td><td><idl< td=""><td>1.26</td><td>90</td><td>639</td></idl<></td></idl<>	33	7.2	1774	211	3.4	6.5	37	5.3	6.8	3.8	26558	<idl< td=""><td>1.26</td><td>90</td><td>639</td></idl<>	1.26	90	639
VIT-1	fibrous fragment	32	<idl< td=""><td>29</td><td>4.6</td><td>2232</td><td>290</td><td>2.1</td><td>1.9</td><td>32</td><td><idl< td=""><td>3.0</td><td>2.1</td><td>6286</td><td><idl< td=""><td>0.73</td><td>61</td><td>127</td></idl<></td></idl<></td></idl<>	29	4.6	2232	290	2.1	1.9	32	<idl< td=""><td>3.0</td><td>2.1</td><td>6286</td><td><idl< td=""><td>0.73</td><td>61</td><td>127</td></idl<></td></idl<>	3.0	2.1	6286	<idl< td=""><td>0.73</td><td>61</td><td>127</td></idl<>	0.73	61	127
55 µm	fibrous fragment	0.38	<idl< td=""><td>21</td><td>3.5</td><td>1112</td><td>293</td><td>0.13</td><td>0.33</td><td>3.8</td><td><idl< td=""><td>0.40</td><td>3.8</td><td>330</td><td><idl< td=""><td>0.11</td><td>3.6</td><td>37</td></idl<></td></idl<></td></idl<>	21	3.5	1112	293	0.13	0.33	3.8	<idl< td=""><td>0.40</td><td>3.8</td><td>330</td><td><idl< td=""><td>0.11</td><td>3.6</td><td>37</td></idl<></td></idl<>	0.40	3.8	330	<idl< td=""><td>0.11</td><td>3.6</td><td>37</td></idl<>	0.11	3.6	37
	fibrous fragment	10	<idl< td=""><td>17</td><td>3.2</td><td>1133</td><td>335</td><td>0.38</td><td>0.48</td><td>13</td><td>1.5</td><td>2.3</td><td>5.0</td><td>1878</td><td><idl< td=""><td>0.46</td><td>51</td><td>153</td></idl<></td></idl<>	17	3.2	1133	335	0.38	0.48	13	1.5	2.3	5.0	1878	<idl< td=""><td>0.46</td><td>51</td><td>153</td></idl<>	0.46	51	153
	fibrous fragment	12	<idl< td=""><td>37</td><td>8.6</td><td>15077</td><td>235</td><td>11</td><td>2.7</td><td>20</td><td>3.2</td><td>8.4</td><td>11.5</td><td>10505</td><td><idl< td=""><td>0.85</td><td>60</td><td>858</td></idl<></td></idl<>	37	8.6	15077	235	11	2.7	20	3.2	8.4	11.5	10505	<idl< td=""><td>0.85</td><td>60</td><td>858</td></idl<>	0.85	60	858
	fibrous fragment	14	<idl< td=""><td>344</td><td>18</td><td>24632</td><td>244</td><td>133</td><td>6.2</td><td>53</td><td>3.4</td><td>24</td><td>11.7</td><td>3268</td><td><idl< td=""><td>0.87</td><td>119</td><td>771</td></idl<></td></idl<>	344	18	24632	244	133	6.2	53	3.4	24	11.7	3268	<idl< td=""><td>0.87</td><td>119</td><td>771</td></idl<>	0.87	119	771
	Average	18		80	7.5	7660	268	25	3.0	27	3.3	7.5	6.3	8137		0.71	64	431
	SD	15		129	5.6	9926	46	53	2.7	18	1.6	8.6	4.2	9718		0.39	39	365
	Median	13		31	6	2003	267	3	2	26	3	5	4	4777		1	60	396
	Maximum	42		344	18	24632	335	133	6.5	53	5.3	24	11.7	26558		1.26	119	858
	Minimum	0.38		17	3.2	1112	211	0.13	0.33	3.8	1.5	0.40	2.1	330		0.11	3.6	37

Tab. 1 LA-ICP-MS spot analyses in molybdenite from W-mineralization hosted by Vítkov topaz greisen (ppm)

\* Instrument detection limits for spot size of 55  $\mu m$ 

molybdenite that originated in connection with opening of inter-lamellar space in molybdenite (Fig. 2b) and also during later stages of molybdenite plastic deformation (Fig. 2c). There is also evidence that tungstenite and other transitional disulfidic W–Mo and Mo–W phases formed after the molybdenite fragmentation (Fig. 2d). Deformation period started after the formation of molybdenite and lasted to the origin of pyrite that is already undeformed and unfractured (Fig. 2e).

The sequence in sulfidic mineral assemblage (from oldest to youngest) at Vítkov was thus as follows:

arsenopyrite  $\rightarrow$  molybdenite  $\rightarrow$  tungstenite  $\rightarrow$  W-Mo/Mo-W disulfides  $\rightarrow$  pyrite

## 4.2. LA-ICP-MS study of molybdenite

The LA-ICP-MS analyses in molybdenite from greisentype tungsten mineralization at Vítkov show variable metal concentrations (Tab. 1). It should be noted that all Ni and Re values were below detection limit while anomalous concentrations of W (up to 26 558 ppm), As (up to 24 632 ppm), Bi (up to 858 ppm), Se (up to 335 ppm), Cu (up to 344 ppm), Zr (up to133 ppm), Ag (up to 53 ppm), Sb (up to 24 ppm), Te (up to 11.7 ppm), Nb (up to 6.5 ppm), Sn (up to 5.3 ppm) and Au (up to 1.26 ppm) were found in a molybdenite grain.

#### 4.3. Electron-microprobe study of major sulfides

Results of electron-microprobe study of relatively homogeneous (inclusions-free parts) molybdenite and tungstenite grains are summarized in Tab. 2 and of arsenopyrite in Tab. 3. The detailed study of chemical composition across the molybdenite grain rimmed by tungstenite (Tab. 2) showed relatively homogeneous Mo (59.60–60.66 wt. %), S (39.83–40.37 wt. %) and W (0.45–0.59 wt. %) values. Rhenium concentrations were below detection limit which varied between 406 and 511 ppm for individual measurements. Tungstenite contains 65.22-69.87 wt. % W, 27.07-27.41 wt. % S and between 3.34 and 5.87 wt. % of Mo with a strong negative correlation between Mo and W (R= –0.997, Fig. 3). Locally, a transition from Mo–W and W–Mo disulfides to tungstenite, which is often in a sharp contact with molybdenite, can be seen (Fig. 4 and 5a–c).

The EMPA of arsenopyrite which is paragenetically the oldest sulfide revealed relatively homogeneous concentrations of As (45.8–48.2 wt. %), Fe (32.9–34.8 wt. %), S (18.2–20.1 wt. %) and variable Co values (0.023–1.3 wt. %), locally with traces of Ni (Tab. 3).



Fig. 3 Mo vs. W plot in tungstenite. Note a significant negative correlation indicating a substitution of  $Mo^{4+}$  for  $W^{4+}$ .

Mineral		wt. %		at. %							
	S	Mo	W	Total	S	Mo	W	Total			
molybdenite	40.26	60.27	0.56	101.10	66.55	33.29	0.16	100			
	40.12	60.49	0.51	101.12	66.40	33.45	0.15	100			
	40.06	60.45	0.49	100.99	66.38	33.48	0.14	100			
	40.29	60.66	0.48	101.42	66.44	33.43	0.14	100			
	40.23	60.00	0.46	100.68	66.65	33.22	0.13	100			
	40.30	60.30	0.45	101.05	66.58	33.29	0.13	100			
	40.28	60.11	0.55	100.94	66.62	33.22	0.16	100			
	40.04	60.02	0.51	100.58	66.53	33.32	0.15	100			
	40.31	60.09	0.57	100.96	66.64	33.20	0.16	100			
	39.93	59.91	0.55	100.39	66.50	33.34	0.16	100			
	40.37	59.90	0.54	100.81	66.75	33.10	0.16	100			
	40.13	59.99	0.53	100.65	66.59	33.26	0.15	100			
	40.10	59.91	0.52	100.53	66.60	33.25	0.15	100			
	39.93	59.64	0.53	100.10	66.61	33.24	0.15	100			
	39.83	59.62	0.59	100.04	66.54	33.28	0.17	100			
	40.26	59.89	0.59	100.74	66.68	33.15	0.17	100			
	40.02	59.80	0.49	100.31	66.60	33.25	0.14	100			
	40.04	60.28	0.58	100.90	66.42	33.42	0.17	100			
	40.18	59.60	0.59	100.36	66.75	33.09	0.17	100			
Minimum	39.83	59.60	0.45		66.38	33.09	0.13				
Maximum	40.37	60.66	0.59		66.75	33.48	0.17				
Average	40.14	60.05	0.53		66.57	33.28	0.15				
tungstenite	27.14	5.77	65.22	98.12	67.11	4.76	28.13	100			
	27.07	3.49	69.31	99.86	67.13	2.89	29.98	100			
	27.20	3.53	69.53	100.26	67.15	2.91	29.93	100			
	27.23	4.47	67.97	99.67	67.10	3.68	29.21	100			
	27.41	3.81	69.65	100.87	67.12	3.12	29.75	100			
	27.34	3.34	69.87	100.55	67.26	2.75	29.99	100			
Minimum	27.07	3.34	65.22		67.10	2.75	28.13				
Maximum	27.41	5.77	69.87		67.26	4.76	29.99				
Average	27.23	4.07	68.59		67.15	3.35	29.50				

Tab. 2 Chemical composition of molybdenite and tungstenite (EMPA)

### 5. Discussion

## 5.1. Tungsten-bearing greisens in the Bohemian Massif

Greisens associated with late Variscan granites (Tischendorf 1989; Štemprok et al. 1997) are a common host of tin and tungsten ores in the Sn–W metallogenic

Tab. 3 Chemical composition of arsenopyrite (EMPA)

(Štemprok and Seltmann 1994). They are metasomatic rocks mostly formed from a granite precursor, and consist of quartz, mica (Li-Fe mica or muscovite; forming either Li-rich or Li-poor greisens), and/or topaz. The greisens are spatially related to the tin and tungsten mineralization of the Younger Intrusive Complex (YIC) granites of the Krušné hory/Erzgebirge granite Batholith. Among the ore minerals, the greisens contain cassiterite, wolframite, and/or scheelite in variable combinations and proportions, in addition to a small amount of sulfides (Štemprok et al. 2005). No tungstenite and/or transitional W-Mo/Mo-W disulfides were reported from these occurrences. The presence of rare micro-inclusions (up to 3 µm across) of tungstenite in cassiterite from granite at Přibyslavice was described by Breiter et al. (2006).

province of Central Europe

The W-bearing greisens and quartz veins are found within a large outer tungsten

metallogenic zone that encircles the granites of the Western Krušné hory/Erzgebirge Batholith (Štemprok and Seltmann 1994), and are characteristic of the western part of the Krušné hory/Erzgebirge ore province. For example, a W-bearing greisen was found at Boží Dar in the eastern contact zone of the Nejdek–Eibenstock Massif by Absolonová (1978) and at Vykmanov by Štemprok (1984). At Boží Dar, the deposits occur

Mineral			wt. %			at.%								
	Fe	Co	Ni	As	S	Total	Fe	Co	Ni	As	S	Total		
arsenopyrite	33.03	1.32	0.12	47.80	18.27	100.55	32.50	1.23	0.11	35.05	31.30	100		
	32.93	1.12	0.15	48.21	18.19	100.60	32.38	1.04	0.14	35.33	31.15	100		
	34.43	0.03	0.05	46.98	19.14	100.62	33.48	0.02	0.04	34.05	32.42	100		
	34.80	0.05	0.01	45.83	20.11	100.81	33.48	0.05	0.01	32.87	33.71	100		
	34.53	0.04	0.03	46.45	19.46	100.51	33.51	0.04	0.03	33.60	32.89	100		
	33.12	1.21	0.10	47.80	18.48	100.71	32.45	1.12	0.10	34.91	31.53	100		
Minimum	32.93	0.03	0.01	45.83	18.19		32.38	0.02	0.01	32.87	31.15			
Maximum	34.80	1.32	0.15	48.21	20.11		33.51	1.23	0.14	35.33	33.71			
Average	33.82	0.64	0.08	47.14	18.99		32.96	0.60	0.07	34.25	32.23			



**Fig. 4** Energy-dispersive X-ray spectroscopy linescan of the distribution of Mo, W and S across the molybdenite-tungstenite boundary. Note a sharp contact between molybdenite and tungstenite which does not indicate any cation exchange.

in the form of quartz-mica and quartz-topaz greisens with wolframite (ferberite), molybdenite, and native bismuth (Absolonová 1978). The W-bearing topaz greisen at Vykmanov contains ferberite rarely with accessory Fe-columbite or is accompanied by scheelite and rare native bismuth (Stemprok et al. 2005). These authors suggested that the greisen formed from CO<sub>2</sub>poor hydrothermal solutions at about 400 °C. The solutions evolved from highly saline brines; these mostly magmatic fluids were taken responsible for the W-Bi specialization of the greisen. Fluid inclusion study in quartz and topaz from W-bearing greisen at Vítkov (Sidorinová and Dobeš in print) identified primary water-rich inclusions that homogenized at 312-358 °C (minimum T of the origin of minerals), however indicating lower salinity fluids (4.0-7.6 wt. % NaCl equivalent). This also differs from other types of tungsten mineralizations bound to a crystalline envelope of the Krkonoše-Jizera Pluton that showed the presence of two-phase  $(H_2O-CO_2)$  inclusions with similar homogenization temperatures and salinities (Vondrovic et al. 2011) or different salinities (Karwowski 1973) when compared to Vítkov.

#### 5.2. Chemistry of molybdenite at Vítkov

It has been suggested that Re is homogeneously distributed in the molybdenite lattice (e.g. Frondel and Wickman 1970). Molybdenite has been shown to be the major host for Re in many hydrothermal ore deposits (McCandless et al. 1993). Rhenium concentrations vary widely in different types of mineral associations with the highest concentrations in porphyry-type Cu–Mo(Au) deposits (Berzina et al. 2005 and others). Molybdenite from topaz greisen at Vítkov yields very low Re but high W concentrations. These low Re contents are consistent with Stein (2006), who reported very low concentrations of Re (sub-parts of ppm) in molybdenite from greisentype associations and suggested that they were indicative of crustal sources of Re.

Drábek et al. (1989 and 1993) first determined Re, W, Se and Te contents in molybdenite separated from different types of Mo-mineralization of the Bohemian Massif. They concluded that molybdenite from greisen-type deposits contains 1–11 ppm Re (avg. 3 ppm, n = 16), 43–9300 ppm W (avg. 450 ppm, n = 21), 15–200 ppm Se (avg. 120 ppm, n = 19) and 2–14.4 ppm Te



(avg. 7 ppm n = 9). Except of Re and Te, such average values are much lower than those of molybdenite from Vítkov (W = 5300 ppm; Tab. 1 or W = 8137 ppm and Se = 268 ppm; Tab. 3).

Pašava et al. (2016) studied *in situ* variation in chemical composition of molybdenite from different types of mineral associations in the Bohemian Massif. The highest average contents of Cu (58 ppm), Zn (45 ppm), As (40 ppm) and Co (3 ppm) were detected in greisenrelated molybdenite while the highest average values of Bi (1150 ppm), Ag (434 ppm), Te (111 ppm), Sb (38 ppm) and Au (24 ppm) were found in molybdenite from gold deposits.

**Fig. 5** Distribution map of Mo, W and S in a sulfide fragment composed of molybdenite, tungstenite and transitional W–Mo and Mo–W disulfidic phases (cf. Fig. 2d).

Similarly to Re, Frondel and Wickman (1970) suggested that tungsten is also often homogeneously incorporated in the molybdenite lattice. The content of W in molybdenite reflects both the abundance of this element in magmatic system and its chemical bond, which depends on the  $S_2/O_2$  activity ratio. The latter strongly controls whether tungsten joins an oxidic or sulfidic bond (e.g. Štemprok 1971; Drábek 1982 and others). We document that W concentrations of up to 0.5 wt. % found in our low-Re inclusion-free molybdenite are very likely related to its presence in the lattice. Higher W concentrations are, however, associated with detected nano- to micron-scale inter-lamellar fillings and/or intergrowths of tungstenite

10 µm



Fig. 6 Representative time-resolved LA-ICP-MS depth profile for molybdenite from Vítkov to show noisiness of signals for Ag, As, Bi, Nb, Sb, Zr and W indicating presence of impurities and contrasting with flat signals for lattice-bound elements (Se and Te).

with molybdenite and also related to transitional W–Mo/ Mo–W disulfides.

The presence of As, Ag, Bi, Nb, Sb, Zr and W impurities and most likely isomorphic Se and Te admixtures are indicated by the differing response from trace elements in the time resolved LA-ICP-MS depth profiles from molybdenite at Vítkov (Fig. 6).

#### 5.3. Tungstenite-molybdenite relationship

Tungstenite (WS<sub>2</sub>) is isotypic with molybdenite (MoS<sub>2</sub>) and the ionic radius of Mo<sup>4+</sup> (0.65 Å) is very close to that of W<sup>4+</sup> (0.66 Å; Shannon 1976). Both these sulfides form hexagonal (2*H*) and trigonal (3*R*) polytypes (e.g., Gait and Mandarino 1970). Moh and Udubasa (1976) studied the phase relations within the Mo–W–S system at 500–1000 °C and reported a complete solid solution between the hexagonal MoS<sub>2</sub> and WS<sub>2</sub>. Similar complete molybdenite–tungstenite solid solution was also described in a low-temperature mineral assemblage from the Felbertal scheelite deposit (Hohe Tauern, Austria)

by Höll and Weber-Diefenbach (1973). Barkov et al. (2000a) reported the molybdenite zoned with respect to tungsten from fenitized megaxenolith in the Khibina alkaline Complex, Kola Peninsula, Russia. Tungsten was strongly enriched in the central zone of molybdenite (up to 5.85 wt. %). We have not confirmed any zonality in the W distribution in molybdenite at Vítkov, though. Štemprok (1971) examined phase relations within the Fe-W-S system and discussed the sulfur fugacity therein. His investigation confirmed the stability of the FeS-W tie line and confirmed the position of WS<sub>2</sub>/W curve above that of FeS/Fe; WS<sub>2</sub>-Fe join was not stable. Conversely, according to Kullerud (1966) and Grover et al. (1975), MoS, is stable with metallic Fe. These differences in behavior indicate different affinities of W and Mo to sulfur (Barkov et al. 2000a, b). At the low  $f(O_2)$ fugacity, elements such as W, Ti or Nb with lithophile behavior can become chalcophile (Barkov et al. 1997; 2000a, b). Noticeable lack of W4+ in molybdenite results from its formation at highly oxidizing conditions (Barkov et al. 2000a).



**Fig. 7** The  $\log fS_2$  vs.  $\log fO_2$  diagram indicating the stability fields of tungstenite, molybdenite, ferberite, scheelite and powellite at 577 °C and 100 MPa fluid pressure. Adapted from Hsu (1976, 1977). Stability fields of iron sulfides and oxides are shown for comparison.

Hsu (1977) studied the stability of molybdenite, scheelite and powellite and Hsu (1976) the stability of tungstenite and ferberite. His results are summarized in a plot of log  $f(S_2)$  vs.  $f(O_2)$  showing the stability fields of tungstenite (WS<sub>2</sub>), ferberite (FeWO<sub>4</sub>), molybdenite (MoS<sub>2</sub>), scheelite (CaWO<sub>4</sub>) and powellite (CaMOO<sub>4</sub>) (Fig. 7). From this diagram follows that tungstenite crystalizes at the low oxygen fugacity while ferberite or scheelite are formed at more oxidizing conditions. Similarly at a higher  $f(O_2)$ , powellite originates instead of molybdenite. The formation of powellite, however, requires over five orders of magnitude higher  $f(O_2)$  than needed for scheelite. Molybdenite is not stable in the presence of hematite (Drábek 1982) and tungstenite is not stable in the presence of iron oxides (Štemprok 1971).

## 5.4. Implications for the origin of sulfide association at Vítkov

Sidorinová and Dobeš (in print) reported only water-rich inclusions from quartz and topaz of greisen at Vítkov. Mineral paragenesis and fluid-inclusion data indicate that oxidizing aqueous fluids rich in rare metals precipitated wolframite and scheelite during oxide-dominated conditions while the change of physico-chemical conditions (strong decrease in  $fO_2/S_2$ ) resulted in the formation of sulfidic assemblage. After arsenopyrite (that scavenged

most of As and Fe available in the system) followed precipitation of molybdenite with up to 0.5 wt. % of W. Later on, after decreasing  $f(O_2)$ , conditions shifted into the stability field of W<sup>4+</sup> ions (below WO<sub>2</sub>–WO<sub>3</sub> buffer; Cygan and Chou 1990). This triggered crystallization of tungstenite in the form of "inclusion-like" lamellae in the inter-lamellar space, parallel to the basal faces of the molybdenite crystals. The contacts between older molybdenite and younger tungstenite are sharp without any exchange of cations (Fig. 4).

Probably still younger were changes in Mo/W ratio in tungstenite growing at outer margins of the molybdenite aggregates. In this case molybdenite–tungstenite contacts remain sharp, but peripheral zones are poorer in W, reflecting most likely later changes in the activity of W<sup>4+</sup> in the mineralizing fluid (Fig. 4 and 5a–c). Zoned tungstenite crystals and sharp tungstenite–molybdenite contacts indicate disequilibrium conditions during their formation.

## 6. Conclusions

The detailed mineralogical investigation of sulfide assemblage from W-bearing topaz greisen at Vítkov followed by LA-ICP-MS and EMP study resulted in the following conclusions:

- 1. The origin of sulfide-rich assemblage started with the precipitation of arsenopyrite, followed by the formation of molybdenite, tungstenite, Mo–W and W–Mo transitional disulfides and, finally, by pyrite. This is for the first time when molybdenite associated with tungstenite and Mo–W and W–Mo disulfidic phases is described from W-bearing topaz greisen in the Bohemian Massif.
- 2. LA-ICP-MS analyses of molybdenite showed extreme concentrations of W (up to 26 558 ppm) accompanied by elevated values of other metals (e.g., As, Bi, Se, Cu, Pb and Ag) and very low Re values (below 0.3 ppm), consistent with the crustal source of metals.
- 3. Electron-microprobe analyses of inclusion-free molybdenite yielded the average value of ~0.5 wt. % of W most likely bound into the crystal lattice. Inclusion-free tungstenite contains on average ~4 wt. % of Mo that shows a strong negative correlation with W, indicating a substitution of Mo<sup>4+</sup> for W<sup>4+</sup>.
- 4. Textural features, along with the stability and phase relationships between molybdenite, tungstenite and transitional Mo- and W-dominated disulfides suggest that precipitation of molybdenite took place at relatively high  $fO_2$ . The subsequent crystallization of tungstenite was triggered by a decrease in  $fO_2$  below the WO<sub>2</sub>-WO<sub>3</sub> buffer. Sharp boundaries between tungstenite and molybdenite and locally zoned tungstenite indicate disequilibrium during their formation.

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